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# Entanglement viscosity and conformal quantum anomaly

based on works:

[G. Prokhorov, 2601.02083 (2026)]

[G. Prokhorov, O. Teryaev, 2604.04222 (2026)]

[D. Lapygin, G. Prokhorov, O. Teryaev, V. Zakharov,  
PRD 112 (2025) 6, 065012]

[R. Khakimov, G. Prokhorov, O. Teryaev, in preparation]

**Quarks-2026,**  
Russia, Petrozavodsk,  
May 18-23, 2026

# Contents

## Introduction

1. **Entanglement viscosity:** *a) universal formulas,*  
*b) unitarity vs irreversibility.*
2. **Entanglement viscosity vs conformal anomaly.**
3. **Novel sum rule** for spectral functions from the isotropy of Unruh radiation:  
*direct verification*
4. **Shear viscosity to entropy density ratio:** *a) local vs global.*  
*b) KSS-bound and causality.*
5. **Bulk viscosity:** *derivation of the bound without holography*
6. **Phenomenological applications**

## Conclusion

**Part 1**

**Introduction**

# Minimal viscosity bounds

Hydrodynamics in linear gradients - corrections to EMT with dissipation:

$$T_{\mu\nu} = T_{\mu\nu}^{\text{ideal}} + T_{\mu\nu}^{\text{diss}}$$

$$T_{\mu\nu}^{\text{ideal}} = (\varepsilon + p)u_\mu u_\nu - pg_{\mu\nu}$$

$$T_{\mu\nu}^{\text{diss}} = -\eta(\nabla_\mu u_\nu + \nabla_\nu u_\mu - u_\mu u^\alpha \nabla_\alpha u_\nu - u_\nu u^\alpha \nabla_\alpha u_\mu) - \left(\zeta - \frac{2}{3}\eta\right) \nabla^\alpha u_\alpha (g_{\mu\nu} - u_\mu u_\nu) + \mathcal{O}(\nabla^2 u)$$

**Bound inspired by string theory:**

- **There are no completely ideal fluids!**
- It is believed that QGP near this limit
- **Does not cover case of Rindler space!**
- Can be derived within membrane paradigm
  
- Another well-known bound for **bulk viscosity** also from holography:  
[Buchel, PLB (2008), arXiv:0708.3459]

$$\frac{\eta}{s} \geq \frac{1}{4\pi}$$

**KSS-bound**



[Kovtun, Son, Starinets, PRL (2005)]

[Maulik Parikh and Frank Wilczek. Phys. Rev. D, 58:064011, 1998.]

$$\frac{\zeta}{\eta} \geq 2 \left( \frac{1}{p} - c_s^2 \right)$$

# Unruh effect

## Formulation:

The **Minkowski vacuum** is perceived by an **accelerated** observer as a medium with a finite **Unruh temperature**

$$T_U = \frac{a}{2\pi}$$

[W. G. Unruh. Notes on black hole evaporation. Phys. Rev., D14:870, 1976.]



[Blasone, (2018), e-Print: 1911.06002]

# Statement of the problem

- Does **Unruh radiation** have **viscosity**? How is it related to the **KSS bound**?
- Direct calculation for **quantum fluid** above the classical membrane.

## First step (known):

[G. Chirco, C. Eling, and S. Liberati. Phys. Rev. D, 82:024010, 2010]

For massless scalar field - **finite shear viscosity**, saturates **KSS bound**

## This work:

We generalize to **“any” theory**:

- demonstrate the role of **Unitarity** and **Causality**
- demonstrate the **relationship** with **irreversibility of renormalization group flows**
- something more ...

[A. B. Zamolodchikov. JETP Lett., 43:730–732, 1986]

## **Part 2**

**Entanglement  
viscosity:  
universal formulas**

# Kubo formulas in curved spacetime

- Linear response theory allows us to write **Kubo formula** in **curved spacetime**:

[Dam T. Son. Acta Phys. Polon. B, 39:3173–3182, 2008]

**Shear viscosity:**

$$\eta(x') = \text{Im} \lim_{\omega \rightarrow 0^+} \frac{i}{\sqrt{g^{00}(x')}\omega} \int_{-\infty}^{\infty} dt \theta(t - t') \cdot e^{i\omega(t-t')} \int_V d^{d-1}x \sqrt{-g(x)} \langle [\hat{T}_{12}(x), \hat{T}_{12}(x')] \rangle$$

**Bulk viscosity:** similar formula with trace of EMT

- Apply Kubo formula to the  **$d$ -dimensional Rindler space**:

$$ds^2 = \rho^2 d\tau^2 - (dx^1)^2 - \dots - (dx^{d-2})^2 - d\rho^2$$

- Consider **Minkowski vacuum state**:

Thermal correlator corresponds to the **Minkowski vacuum** one (with a change of coordinates)

[W. G. Unruh and N. Weiss. Phys. Rev., D29:1656, 1984]

# Universal spectral representation

- Universal spectral representation:**

[A. Cappelli, D. Friedan, and J. I. Latorre. Nucl. Phys. B, 352:616–670, 1991]

[M. Smolkin and S. N. Solodukhin, Phys. Rev. D 91, no.4, 044008 (2015)]

[S. N. Solodukhin. The a-theorem and entanglement entropy. 4 2013.]

$$\langle 0 | \hat{T}_{\alpha\beta}(x) \hat{T}_{\rho\sigma}(x') | 0 \rangle_M = \frac{A_d}{(d-1)^2} \int_0^\infty d\mu c^{(0)}(\mu) \Pi_{\alpha\beta,\rho\sigma}^{(0)}(\partial') G_d(x-x', \mu) + \frac{A_d}{(d-1)^2} \int_0^\infty d\mu c^{(2)}(\mu) \Pi_{\alpha\beta,\rho\sigma}^{(2)}(\partial') G_d(x-x', \mu)$$

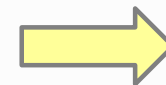
**spectral function**

scalar propagator

- Valid for “any” theory: different spins, conformal, nonconformal, interacting...

- For unitary theory:  $c^{(0)}(\mu) \geq 0$   $c^{(2)}(\mu) \geq 0$

- Only  $c^{(0)}(\mu)$  for 2d theory: positivity  $c^{(0)}(\mu) \geq 0$



irreversibility of renormalisation group flows

**Unitarity provides irreversibility of renormalisation group flows**

[A. B. Zamolodchikov. JETP Lett., 43:730–732, 1986]

# Key integral: problem

- Use universal **spectral representaion**
- Use equation for (Whightman) scalar Green function:  $\square G_d(x - x') = -\mu^2 G_d(x - x')$
- Use **descent method**:  $\int dx^{d-2} G^d(x) \rightarrow G^2(x)$

**Main integral  
to be found:**

$$\eta(\rho') = k_d \rho' \int d\mu \mu^4 c^{(2)}(\mu) \operatorname{Im} \lim_{\omega \rightarrow 0^+} \frac{i}{\omega} \int_0^\infty d\tau e^{i\omega\tau} \\ \cdot \int_0^\infty d\rho \frac{\rho}{2\pi} \left\{ K_0 \left( \mu \sqrt{\rho^2 + \rho'^2 - 2\rho\rho' \cosh(\tau - i\varepsilon)} \right) - (\tau \rightarrow -\tau) \right\}$$

- The central question is **how to find this integral?**
  - Includes non-tabular integrals
  - Includes special divergent functions with a cut
  - Up to our knowledge, there is no description in the literature on how to find them.



# Key integral: solution

- However, everything **becomes easy** in the **momentum representation!**



$$\frac{1}{2\pi} K_0 \left( \mu \sqrt{\rho^2 + \rho'^2 - 2\rho\rho' \cosh(\tau - i\varepsilon)} \right) =$$

$$= \frac{1}{\pi^2} \int_0^\infty d\lambda K_{i\lambda}(\mu\rho) K_{i\lambda}(\mu\rho') \cosh(\pi\lambda - i\lambda(\tau - i\varepsilon))$$

[A P Prudnikov, Yu A Brychkov, and O I Marichev. Integrals and series: Special functions vol 2.]

Dependence on  $\rho$  and  $\tau$  is **factorized**

- Now integral over the distance from the horizon can be easily found:

$$\int_0^\infty d\rho \rho K_{i\lambda}(\mu\rho) = \frac{\pi\lambda}{2\mu^2 \sinh(\pi\lambda/2)}$$

- Integral over  $\tau$  leads to frequency parts of delta function – principal value part are real and drops out

$$\eta = \frac{k_d \rho'}{4} \lim_{\omega \rightarrow 0^+} \frac{1}{\omega} \int_0^\infty d\mu c^{(2)}(\mu) \mu^2 \int_0^\infty d\lambda \frac{\lambda K_{i\lambda}(\mu\rho')}{\sinh(\pi\lambda/2)}$$

$$\cdot \left\{ e^{\pi\lambda} \delta(\omega - \lambda) + e^{-\pi\lambda} \delta(\omega + \lambda) - e^{\pi\lambda} \delta(\omega + \lambda) - e^{-\pi\lambda} \delta(\omega - \lambda) \right\}$$

# Result: final universal formulas

- **Entanglement** (local) **shear viscosity** of **Unruh** radiation:

$$\eta(\rho) = k_d \rho \int_0^\infty d\mu c^{(2)}(\mu) \mu^2 K_0(\mu\rho)$$

The calculation for **bulk viscosity** is very similar:

$$\zeta(\rho) = \frac{2k_d \rho}{(d-1)^2} \int_0^\infty d\mu c^{(0)}(\mu) \mu^2 K_0(\mu\rho)$$

- Valid in general case (conformal, nonconformal, interacting...).
- $c^{(2)}(\mu)$  defines **shear viscosity**.
- $c^{(0)}(\mu)$  defines **bulk viscosity**.
- In accordance with particular direct calculations (scalar, Dirac, photon):  
[G. Chirco, C. Eling, and S. Liberati. Phys. Rev. D, 82:024010, 2010]  
[D. Lapygin, G. Prokhorov, O. Teryaev, V. Zakharov, Phys.Rev.D 112 (2025) 6, 065012]

# Discussion: Unitarity vs Irreversibility

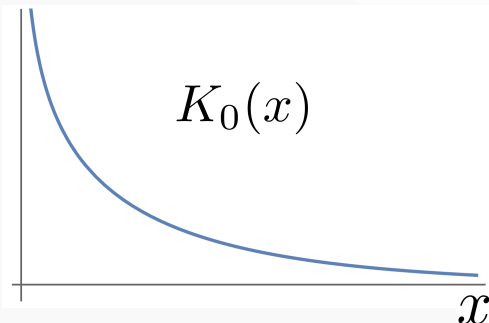
$$\eta(\rho) = k_d \rho \int_0^\infty d\mu c^{(2)}(\mu) \mu^2 K_0(\mu\rho)$$
$$\zeta(\rho) = \frac{2k_d \rho}{(d-1)^2} \int_0^\infty d\mu c^{(0)}(\mu) \mu^2 K_0(\mu\rho)$$

- **Unitarity:**

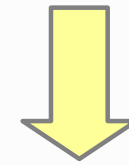
$$c^{(0)}(\mu) \geq 0 \quad c^{(2)}(\mu) \geq 0$$

- Include modified Bessel function of the second kind:

$$K_0(x) > 0$$



$$\eta, \zeta \geq 0$$



$$\partial_\mu s^\mu \geq 0$$

- **Unitarity** underlies **thermal irreversibility** for a subsystem separated by a horizon
- **Analogy** with the **irreversibility** of **renormalization group flows** in 2d.

[Andrea Cappelli, Daniel Friedan, and Jose I. Latorre. Nucl. Phys. B, 352:616–670, 1991]

# **Part 3**

**Entanglement  
viscosity and  
conformal anomaly**

# New area: anomalous transport effects

New (**non-dissipative**) effects are predicted at the intersection of quantum field theory and gravity (*only some of them*):

- Chiral Magnetic Effect (**CME**):

[K. Fukushima, D. E. Kharzeev, H. J. Warringa, PRD 78, 074033 (2008), 0808.3382]

$$\langle \hat{j}^\mu \rangle = C e^2 \mu_A B^\mu$$

Associated with **axial anomaly** in the electromagnetic field:

$$\langle \partial_\mu \hat{j}_A^\mu \rangle = -\frac{C e^2}{8} \epsilon^{\mu\nu\alpha\beta} F_{\mu\nu} F_{\alpha\beta}$$

- Axial Vortical Effect (**AVE**):

[D. T. Son, P. Surowka, PRL 103, 191601 (2009), 0906.5044]

$$\langle \hat{j}_A^\mu \rangle = C(\mu^2 + \mu_A^2) \omega^\mu$$

Associated with an **axial anomaly** in the **gravitational field**:

$$\langle \nabla_\mu \hat{j}_A^\mu \rangle = \mathcal{N} \epsilon^{\mu\nu\alpha\beta} R_{\mu\nu\rho\sigma} R_{\alpha\beta}{}^{\rho\sigma}$$

- Thermal part of AVE:

[K. Landsteiner, E. Megias, F. Pena-Benitez, PRL 107, 021601 (2011), 1103.5006]

$$\langle \hat{j}_A^\mu \rangle \sim \mathcal{N} T^2 \omega^\mu$$

- Kinematic Vortical Effect (KVE)**

[G. Y. Prokhorov, O. V. Teryaev, V. I. Zakharov, PRL 129, 151601 (2022), 2207.04449]

$$\langle \hat{j}_\mu^A \rangle = (\lambda_1 \omega^2 + \lambda_2 a^2) \omega_\mu$$

**Connection with anomaly:**

$$\lambda_1 - \lambda_2 = 32\mathcal{N}$$

# Entanglement viscosity and conformal anomaly

- **Conformal anomaly** effects are **not so well studied**.

[M. N. Chernodub, Phys. Rev. Lett. 117, 141601 (2016), arXiv:1603.07993 [hep-th]]

[C.-S. Chu and R.-X. Miao, Phys. Rev. Lett. 121, 251602 (2018), arXiv:1803.03068 [hep-th]]

[Yang, Gao, Liang, Prokhorov, Shi Pu, Teryaev, Zakharov, in preparation]

- Conformal gravitational anomaly in 4d:

$$\langle \hat{T}_{\mu}^{\mu} \rangle = a(-H + \frac{2}{3}\square R) - bE_4 + c\square R$$

$$\eta = 8a\alpha^3$$

- Entanglement viscosity – **novel anomalous effect** in systems with **extreme acceleration**  $\alpha$ :

- “Cheshire cat’s grin” of conformal gravitational anomaly in flat spacetime:



# **Part 4**

**Novel sum rule  
for spectral  
densities**

# Modular Hamiltonian perturbation theory

- There is a technique to find perturbatively effects of small angular deficit using modular Hamiltonian (which is actually a boost operator):

[M. Smolkin and S. N. Solodukhin, Phys. Rev. D 91, no.4, 044008 (2015)]

[V. Rosenhaus and M. Smolkin. JHEP, 12:179, 2014.]

$$\lim_{\nu \rightarrow 1} \frac{\partial}{\partial \nu} \langle \hat{T}_{ij} \rangle = -\langle \hat{T}_{ij} \hat{K}_0 \rangle$$

- Euclidean Rindler space - describes an accelerated system at a finite temperature

$$ds^2 = \rho^2 d\tau^2 + (dx^1)^2 + \dots + (dx^{d-2})^2 + d\rho^2$$

*It is also a conical space*

- The derivative with respect to the angular deficit corresponds to the derivative with respect to temperature:

$$\nu = \frac{2\pi T}{a} \quad \longrightarrow \quad \lim_{\nu \rightarrow 1} \frac{\partial}{\partial \nu} \quad \Longleftrightarrow \quad \lim_{T \rightarrow T_U} \frac{\partial}{\partial T}$$

# Anisotropy of pressure: Modular Hamiltonian

- Using the results of [Smolkin&Solodukhin](#), one can write the derivatives of pressure with respect to temperature:

$$\langle \hat{T}_\nu^\mu \rangle(a, T) = \text{diag}(\varepsilon, -p_\perp, \dots, -p_{||})$$

$$\left. \frac{\partial p_\perp}{\partial T} \right|_a = \frac{-2\pi\rho A_d}{(d-1)^2\Gamma(d)} \int_0^\infty d\mu \mu^2 K_0(\mu\rho) \cdot \left\{ c^{(0)}(\mu) - (d-1)c^{(2)}(\mu) \right\},$$

$$\left. \frac{\partial p_{||}}{\partial T} \right|_a = \frac{2\pi A_d}{(d-1)^2\Gamma(d)} \int_0^\infty d\mu \mu K_1(\mu\rho) \cdot \left\{ c^{(0)}(\mu) + (d-1)(d-2)c^{(2)}(\mu) \right\}$$

- The key point – **different** formulas for **longitudinal** and **transverse pressures**:

**In general, pressure can be anisotropic!**

# Pascal's law

- Let us require (or introduce a criterion) **isotropy** of **Unruh radiation**:

$$\frac{\partial p_{||}}{\partial T} = \frac{\partial p_{\perp}}{\partial T}$$



- Novel sum rule**, relating  $c^{(0)}(\mu)$  and  $c^{(2)}(\mu)$  spectral functions:

$$\int_0^{\infty} d\mu \left\{ c^{(0)}(\mu) \mathcal{A}^{(0)}(\mu, \rho) + c^{(2)}(\mu) \mathcal{A}^{(2)}(\mu, \rho) \right\} = 0$$

where:

$$\mathcal{A}^{(2)}(\mu, \rho) = -(d-1)\mu^2 K_0(\mu\rho) + (d-1)(d-2)\frac{\mu}{\rho} K_1(\mu\rho)$$
$$\mathcal{A}^{(0)}(\mu, \rho) = \mu^2 K_0(\mu\rho) + \frac{\mu}{\rho} K_1(\mu\rho)$$

- Postulation of **Pascal law** for Unruh radiation
- It is a criterion of **fluidity** (anisotropic pressure in solid body)
- Can be compared with the **Burkhardt-Cottingham sum rule**

[Hugh Burkhardt and W. N. Cottingham. *Annals Phys.*, 56:453–463, 1970]

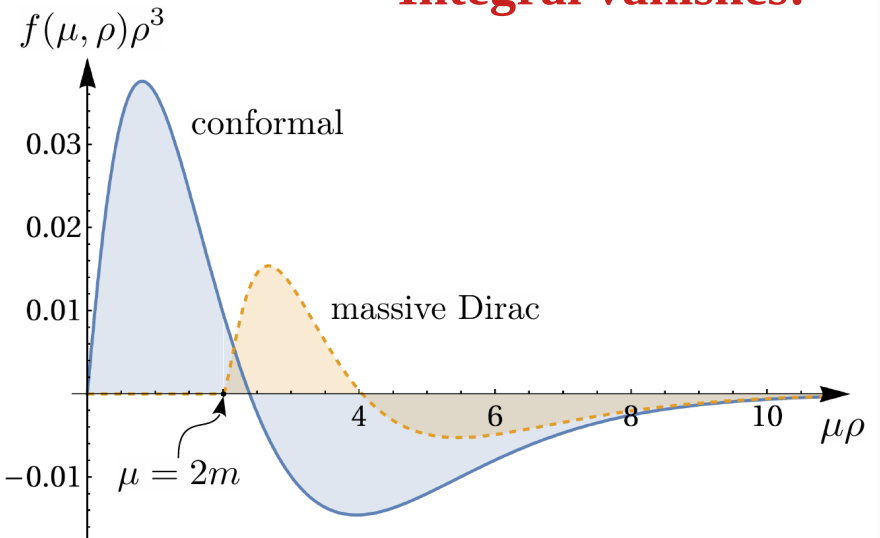
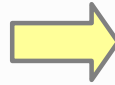
# Direct verification

- Can be **verified** in **conformal** case in **any dimensions**.
- Can be **verified** in **nonconformal** case: free massive Dirac field in any dimensions.

Spectral functions of **massive fermions**:

$$c^{(2)}(\mu, m) = \frac{2^{\lfloor d/2 \rfloor} (d-1) \Gamma\left(\frac{d}{2}\right)^2}{8\pi^d} \mu^{d-3} \left(1 - \frac{4m^2}{\mu^2}\right)^{\frac{d-1}{2}} \cdot \left(1 + \frac{8}{d-1} \cdot \frac{m^2}{\mu^2}\right) \theta(\mu - 2m).$$

$$c^{(0)}(\mu, m) = \frac{2^{\lfloor d/2 \rfloor} (d^2 - 1) \Gamma\left(\frac{d}{2}\right)^2}{2\pi^d} m^2 \mu^{d-5} \left(1 - \frac{4m^2}{\mu^2}\right)^{\frac{d-1}{2}} \cdot \theta(\mu - 2m),$$



[Andrea Cappelli, Daniel Friedan, and Jose I. Latorre. Nucl. Phys. B, 352:616–670, 1991]

## Scalar field:

- Satisfied for conformal scalar field
- Violated for massless nonconformal scalar field
- Open question - improved massive scalar field

- Anisotropic contribution

$$T_{\mu\nu} \sim a_\mu a_\nu$$

[Frolov, V., & Serebryanyi, E. (1987), Phys. Rev. D, 35, 3779–3782]

- Surface terms in the modular Hamiltonian can be important

[Herzog, C., & Nishioka, T. (2016). JHEP, 12, 138.]

# **Part 5**

**Viscosity to  
entropy density  
ratio:  
local vs global**

# Viscosity to entropy density ratio: local vs global

- What is about **KSS bound** for entanglement viscosity?
- Central universal expression for entanglement shear viscosity:

$$\eta(\rho) = k_d \rho \int_0^\infty d\mu c^{(2)}(\mu) \mu^2 K_0(\mu\rho)$$

- **Heat capacity** from is the **temperature derivative** of **energy density**:

- can be obtained via modular Hamiltonian perturbative approach

[M. Smolkin and S. N. Solodukhin, Phys. Rev. D 91, no.4, 044008 (2015)]

$$\left. \frac{\partial \varepsilon}{\partial T} \right|_a = \frac{2\pi\rho A_d}{(d-1)^2 \Gamma(d)} \int_0^\infty d\mu \left[ \mu^2 K_0(\mu\rho) + \frac{\mu}{\rho} K_1(\mu\rho) \right] \left\{ c^{(0)}(\mu) + (d-1)(d-2)c^{(2)}(\mu) \right\}$$



use **isotropy sum rule**

$$c_V(\rho) = \frac{2\pi\rho A_d}{\Gamma(d)} \int_0^\infty d\mu c^{(2)}(\mu) \mu^2 K_0(\mu\rho)$$

only via  $c^{(2)}(\mu)$

- **One can easily see:**

$$\frac{\eta}{c_V} = \frac{1}{4\pi}$$

- Valid in general case (conformal, nonconformal, interacting...)
- Obtained without holography

# Viscosity to entropy density ratio: local vs global

- Can be transformed to the form with entropy density
- Use relation:

$$c_s^2(a, T) = \left. \frac{\partial p}{\partial \varepsilon} \right|_a = \frac{\partial p(a, T)/\partial T}{\partial \varepsilon(a, T)/\partial T} = \frac{s}{c_V}$$

we obtain:

$$\frac{\eta}{c_V} = \frac{1}{4\pi}$$




$$\frac{\eta}{s} = \frac{1}{4\pi c_s^2}$$

$$\frac{\eta}{s} > \frac{16}{25 \cdot 4\pi}$$

Compare with:

[Mauro Brigante, Hong Liu, Robert C. Myers, Stephen Shenker, and Sho Yaida. Phys. Rev. Lett., 100:191601, 2008]

- For example, in conformal limit  $c_s^2 = \frac{1}{3}$  then  $\frac{\eta}{s} = \frac{3}{4\pi}$
- Relationship between **KSS bound** and **causality**:

**Causality:**  $c_s^2 < 1$    $\frac{\eta}{s} > \frac{1}{4\pi}$

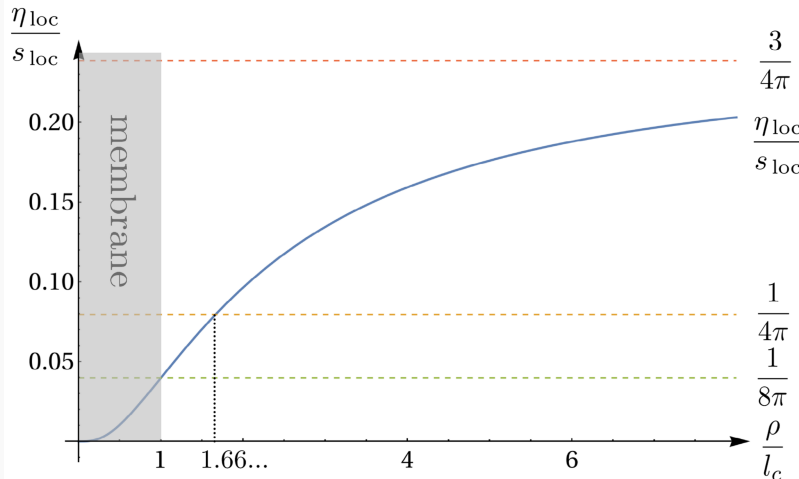
# Viscosity to entropy density ratio: local vs global

- There are **local** and **global** quantities:

## Local

- At certain distance  $\rho$  from the horizon
- We can now regularize integrals by stretched horizon  $\rho > l_c$
- Any CFT in 4d:

$$\frac{\eta}{s} = \frac{1}{4\pi c_s^2}$$



$$\frac{\eta_{loc}}{s_{loc}}(\rho) = f(\rho/l_c) = \frac{3\rho^4 \left[ \rho^4 + 4\rho^2 l_c^2 - 5l_c^4 - 4l_c^2(2\rho^2 + l_c^2) \ln\left(\frac{\rho}{l_c}\right) \right]}{4\pi(\rho^2 - l_c^2)^4}$$

- Universal function** for any CFT in 4d

## Global

$$\eta_{glob} = \int_{l_c}^{\infty} \eta(\rho) d\rho = \frac{C_T \pi^2}{480 l_c^2}$$

$$s_{glob} = \int_{l_c}^{\infty} s(\rho) d\rho = \frac{C_T \pi^3}{120 l_c^2}$$

$$\frac{\eta_{glob}}{s_{glob}} = \frac{1}{4\pi}$$

- KSS bound is saturated globally** for any CFT in 4d

(all the previous discussion – for  $l_c \rightarrow 0$ )

# “Wandering” Planck constant

- In the original holographic derivation (or within membrane paradigm) of KSS bound the **viscosity** is “**classical**” – Planck constant comes from the “**quantum**” Bekenstein-Hawking **entropy**:

$$\eta \sim \mathcal{O}(\hbar^0)$$

$$s \sim \frac{A}{\hbar G} \sim \mathcal{O}(\hbar^{-1})$$

$$\frac{\eta}{s} = \frac{\hbar}{4\pi}$$

- In our case, the **viscosity** is **quantum**: determined by a **one-loop** diagram, contains Planck constant:

$$\eta_{\text{loc}} \sim \frac{T_U^3}{\hbar^2}$$

$$s_{\text{loc}} \sim \frac{T_U^3}{\hbar^3}$$

$$\frac{\eta}{s} = \frac{\hbar}{4\pi}$$

However, the result is the same:  
“**Wandering**” **Planck constant**

# Higher spins

- **Negative viscosity** for **higher spins** (without supergravity)?

[R. Khakimov, G. Prokhorov, O. Teryaev, in preparation]

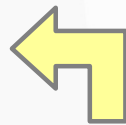
- **Rarita-Schwinger-Adler** theory

$$\mathcal{L} = -\varepsilon^{\alpha\beta\mu\nu} \bar{\psi}_\alpha \gamma_5 \gamma_\mu \partial_\nu \psi_\beta + i\bar{\lambda} \gamma^\mu \partial_\mu \lambda - im\bar{\lambda} \gamma^\mu \psi_\mu + im\bar{\psi}_\mu \gamma^\mu \lambda$$

[Stephen L. Adler. Phys. Rev. D, 97(4):045014, 2018]

- We show that it has features of **conformal theory** on the one-loop level in the limit:  $m \rightarrow \infty$  (unlike usual Rarita-Schwinger field)

- **Negative viscosity**



**Negative conformal central charge**

**Entropy density** is also **negative**

**KSS-bound** is **saturated**

$$\frac{\eta_{glob}}{s_{glob}} = \frac{1}{4\pi}$$

**Part 5**

**Bulk viscosity**

# Viscosity to entropy density ratio: local vs global

- What is about **entanglement bulk viscosity**?

$$\zeta(\rho) = \frac{2k_d\rho}{(d-1)^2} \int_0^\infty d\mu c^{(0)}(\mu)\mu^2 K_0(\mu\rho)$$

- Use **isotropy sum rule**:

-- entropy density

-- heat capacity

Speed of sound

$$c_s^2(a, T) = \frac{s}{c_V}$$

- Use similar “spectral” formula for shear viscosity
- Everything is expressed through integrals with  $c^{(0)}(\mu)$  and  $c^{(2)}(\mu)$

Simple algebra

- **Bulk viscosity** saturates **known bound**:

$$\frac{\zeta}{\eta} = 2 \left( \frac{1}{d-1} - c_s^2 \right)$$

[Buchel, PLB (2008),  
arXiv:0708.3459]

- Obtained **without holography**
- Valid in general case (conformal, nonconformal, interacting...)

**Part 6**

**Phenomenological  
applications**

# Phenomenological applications

- Are there any **phenomenological applications**?
- **New anomalous effect** in systems with **extreme acceleration**:

$$\eta, \zeta \sim a^3$$

extreme acceleration in HICs:

[G. Yu. Prokhorov, D. A. Shohonov, O. V. Teryaev, N. S. Tsegelnik, and V. I. Zakharov. Phys. Rev. C, 112(6):064907, 2025.]

- **Local ratio** with speed of sound:

expressed in terms of general hydrodynamic quantities



can be more general



**compare with QGP**

QGP viscosity from the experiments with HICs:  $0.05 < \eta/s < 0.2$

[John W. Harris and Berndt Müller. "QGP Signatures" Revisited. Eur. Phys. J. C, 84(3):247, 2024]

**Naive prediction** of our formula:

$$\frac{\eta}{s} = \frac{1}{4\pi c_s^2}$$



$$c_s^2 \sim 1/3$$



$$\eta/s \sim 0.24$$

**Look on the Lattice!**

**Part 7**

**Conclusion**


# Key Results

- Thermal quantum radiation in a space with a horizon has an entanglement viscosity.
- We have derived universal formulas for entanglement viscosities via spectral densities.
- We have demonstrated a direct relationship between thermodynamic irreversibility and unitarity in this case. This is a direct analogy with the irreversibility of renormalization group flows.
- We have shown that the Unruh radiation isotropy conditions lead to a new sum rule relating different fundamental spectral functions.
- We have explicitly tested this sum rule for conformal and nonconformal theory (massive fermions) in any number of dimensions.
- Without using holography, we showed that globally for the conformal theory the KSS bound is saturated.
- However, locally the ratio contains the speed of sound.
- This demonstrates the relationship between the KSS bound and causality.
- Without using holography, we showed that the bulk entanglement viscosity of saturates another known bound.
- We have shown that the entanglement shear viscosity is a new anomalous transport effect associated with the conformal gravitational anomaly.
- The obtained relationships with the speed of sound are of interest from a phenomenological point of view: QCD on Lattice?

# Summary & Conclusion

- **Universal formulas** for entanglement shear and bulk viscosities in terms of fundamental **spectral functions** are obtained:
  - Valid for conformal and nonconformal case
  - Demonstrate relationship: **Thermal irreversibility**  $\iff$  **QFT unitarity**
  - Analogy with irreversibility of **renormalization group flows**
  - In accordance with direct calculations
- **Novel sum rule** for different spectral functions is obtained from **Pascal law** for Unruh radiation:
  - Directly verified for **CFT** in d dimensions
  - Directly verified for **nonconformal** case (massive fermions) in d dimensions
- It is shown (**without using holography**) that the entanglement viscosity **saturates known bounds**:
  - The **KSS bound** is **globally saturated**
  - The bound for bulk viscosity is saturated
  - **Novel** formula for **local ratio** with speed of sound:

$$\frac{\eta}{s} = \frac{1}{4\pi c_s^2}$$


$$k_d = \frac{\pi^{d/2}}{(d+1)2^{d-1}\Gamma(d)\Gamma(d/2)}$$