# Limits on ultra-high energy photons flux from Telescope Array's data





## 1. Neural networks

### The idea of machine learning



"Program" is a fixed, created by a human algorithm of solving the problem.

Manually created algorithms

"Programm" - learned on the examples optimal algorithm.

Programs finding optimal algorithms

#### An illustration



## 2. Telescope Array





## Telescope Array is the largest cosmic ray observatory in the Northern hemisphere

(700 km<sup>2</sup>, 507 surface + 3 fluorescent detectors)





## **Motivation**

Detection of EeV photon-induced air showers would indicate new physics

To search for such events, it is desirable to have as big exposure time as possible

- Surface detectors operate almost at all times.
- To get the best separation of photon- and hadron-induced air showers, employ Neural Networks.



hadrons

### Approach

- 1. Train a neural network to distinguish photon- and hadron-induced air-showers.
  - Use Monte-Carlo simulated photon- and proton-induced events for training. This provides conservative limits for photon search.
- 2. Introduce classification threshold.
  - Optimize by requiring best sensitivity to photon candidates in case of the null hypothesis, i.e., no photons in the data.
- 3. Verify applicability to real data analysis
- 4. Set limits by analyzing the data

### Neural network

#### Neural network's blocks:

- Spatial detectors bundle (geometrical features)
- Temporal detector bundle (overall information)
- Reconstruction parameters (high-level information)



## **Detectors bundle**

- x, y, and z coordinates of the detector
- Detector's total signal
- Time of the plane front arrival
- Difference in time between the start of the recorded signal and the wavefront arrival
- Masks:
  - Was triggered?
  - Was saturated?
  - Was excluded from the geometry fit?





## Features from waveforms

- 128 bins (20ns each) of the signal in upper and lower detectors
- Real-time calibration tables, signal measured in "Minimal ionization particle"



## Temporal detector bundle



#### Dataset

- Full MC simulation of 14 years of TA operation time
- Employ 4 high energy hadronic interaction models: QGSJET-II-03, 04, SIBYLL, EPOS (reduce systematic error).
- Equal-efficiency mixture of protonand photon-induced events (no unphysical energy related bias).



#### Model dependence

1) Select one of the hadronic models;

2) Do not use it for training;

3) Compare neural network predictions for the hidden and other hadronic models.

To analyze experimental data, we employ neural network trained on all of the hadronic interaction models.



## **Training tricks**

- 1. Use a special loss function, *focal loss,* forcing neural network to pay more attention to events that are hard to distinguish.
- 2. Trained an ensemble of neural networks and averaged their predictions. This significantly reduces false-positive rate.



3. Fine-train neural network on experimental data that was classified as protons with high confidence.

## Aplicability for analyzing experimental data



## Finding optimal classification threshold



#### Results



Energy, log10 scale, eV	Number of photon candidates
19	0
19.5	0
20	1

# Thank you for attention!

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