

Entanglement Entropy In De Sitter Spacetime: No Pure States For Conformal Matter

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Big Picture Motivation

- ▶ Our universe approaches **de Sitter spacetime**.
- ▶ For a static observer, de Sitter spacetime contains **causally disconnected regions** separated by **cosmological horizon**.
- ▶ Cosmological and black hole horizons share similar properties: **temperature, entropy, Hawking radiation**.
- ▶ However, cosmological horizon is **observer-dependent**, in contrast to black hole geometry. Microscopic interpretation of the cosmological horizon is not clear.

Big Picture Motivation

The microscopics of cosmological horizon thermodynamics could potentially rely on **holography**. For example, the gravitational fine-grained entropy (for AdS black hole) was first derived via $\text{AdS}_{d+1}/\text{CFT}_d$ correspondence (Ryu-Takayanagi formula). However, the de Sitter holography is still a subject to debate (known examples are dS/CFT correspondence, different versions of static patch holography — worldline holography, stretched horizon holography).

Big Picture Motivation

- ▶ Much of a progress has been achieved in deriving gravitational fine-grained entropy in black hole backgrounds by performing calculations using **semi-classical gravitational path integral**.

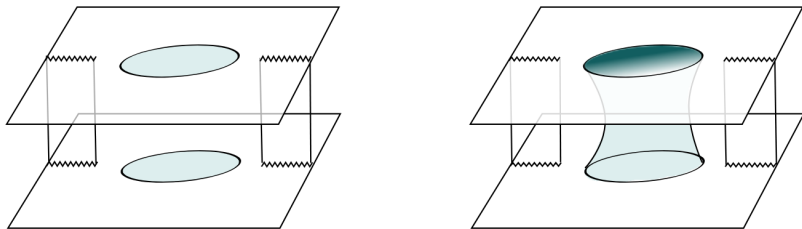


Рис.: Two different saddles in $n = 2$ Euclidean gravitational path integral in the presence of gravity in the shaded region. On the left is the Hawking saddle, on the right — the replica wormhole where gravity dynamically glues together the shaded regions. Replica wormholes reproduce the **island rule**.

Big Picture Motivation

- ▶ Non-trivial saddle points in the path integral (the so-called **replica wormholes**) define the gravitational von Neumann entropy (**QES/island rule**).
- ▶ The island rule is a **non-perturbative** effect.
- ▶ The original derivation of the island rule for AdS_2 black hole relies on the fact that the full manifold $\mathcal{M} \simeq \text{AdS}_2 \cup M_2$. The AdS_2 is dynamical gravity, while Minkowski bath M_2 realizes the **asymptotically flat** b.c. The “radiation” (CFT_2) is collected in M_2 . This allows to consider a replica manifold and use the replica trick.

Big Picture Motivation

- ▶ In dS, event horizon lies between the static observer and the null infinity.
- ▶ Therefore, there is no unambiguous way to couple the thermal bath in dS.
- ▶ Even if the bath is coupled to asymptotic infinity, the Euclidean gravitational path integral will not obtain the corresponding contribution, since only the static patch survives in Euclidean signature.
- ▶ Therefore, for dS the problem does not have a solution neither from holography nor from replica calculation.

Entropy In The Presence Of Gravity

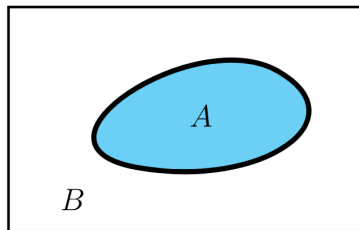
The “island rule” for entanglement entropy of QFT in gravitational systems reads

$$S(R) \simeq \min_{\mathcal{I}} \left\{ \text{ext}_{\mathcal{I}} \left[\frac{\text{Area}(\partial\mathcal{I})}{4G_N} + S_m(R \cup \mathcal{I}) \right] \right\}, \quad (1)$$

where

- ▶ $\Sigma \simeq R \cup \bar{R}$ is the Cauchy surface, on which the state of QFT is defined;
- ▶ R is the region of interest, the entropy of which we aim to calculate;
- ▶ \mathcal{I} — the “island” defined by extremization, $\partial\mathcal{I}$ — its boundary;
- ▶ S_m — entanglement entropy of QFT on the fixed classical background.

Purity And Complementarity Of Entanglement Entropy



- ▶ Let the total system ρ_{tot} consist of two subsystems: $A \cup B$. The Hilbert space of the total system \mathcal{H}_{tot} is given by

$$\mathcal{H}_{tot} \simeq \mathcal{H}_A \otimes \mathcal{H}_B. \quad (2)$$

- ▶ The reduced density matrices of the subsystems are defined as

$$\rho_A = \text{Tr}_{\mathcal{H}_B} \rho_{tot}, \quad \rho_B = \text{Tr}_{\mathcal{H}_A} \rho_{tot}. \quad (3)$$

Purity And Complementarity Of Entanglement Entropy

- ▶ Entanglement entropy of a subsystem X is defined as

$$S_X = -\text{Tr} \rho_X \log \rho_X, \quad X = A, B. \quad (4)$$

- ▶ Basic properties of entanglement entropy of interest:
 - ▶ **Purity**: if the system X is in a pure state, then $\rho_X^2 = \rho_X$, therefore, $S_X = 0$.
 - ▶ **Complementarity**: if the total state $X \cup \bar{X}$ is pure, then entanglement entropies of the subsystem X and its complement \bar{X} are the same: $S_X = S_{\bar{X}}$.
 - ▶ **Araki-Lieb triangle inequality**: $|S_A - S_B| \leq S_{A \cup B}$.
 - ▶ $S_X \leq S_{\text{GH}}(X) \propto \text{Area}(X)$.

Setup

The metric of dS_4 is given by

$$ds^2 = -f(r)dt^2 + \frac{dr^2}{f(r)} + r^2 d\Omega_2^2, \quad f(r) = 1 - \frac{r^2}{\ell^2} \quad (5)$$

The radial distance $d^2(\mathbf{x}, \mathbf{y})$ between two points can be calculated as

$$d^2(\mathbf{x}, \mathbf{y}) = \frac{2\sqrt{f(x)f(y)}}{\kappa_c^2} [\cosh \kappa_c(r_*(x) - r_*(y)) - \cosh \kappa_c(t_x - t_y)] \quad (6)$$

where $\kappa_c = 1/\ell$ is surface gravity, and the tortoise coordinate $r_*(r)$ is

$$r_*(r) = \frac{\ell}{2} \log \frac{\ell + r}{|\ell - r|} = \begin{cases} \ell \operatorname{arctanh} r/\ell, & r < \ell, \\ \ell \operatorname{arctanh} \ell/r, & r > \ell. \end{cases} \quad (7)$$

Setup

- ▶ We use **partial reduction** from $d = 4$ to $d = 2$.
- ▶ We add matter represented by c **free massless Dirac fermions**:

$$\mathcal{S}_{\text{pure } dS_4} \rightarrow \mathcal{S}_{\text{partially reduced } dS_2} + \mathcal{S}_{\text{CFT}_2 \text{ fermions}}.$$

- ▶ Entropy of conformal matter for one interval

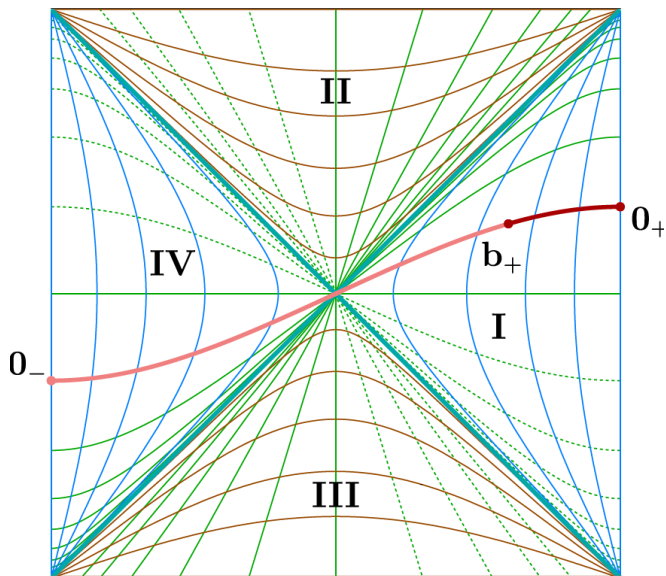
$$S_m = \frac{c}{6} \log \frac{d^2(\mathbf{x}, \mathbf{y})}{\epsilon^2}. \quad (8)$$

- ▶ Entropy of N intervals is given by

$$S_m = \frac{c}{6} \sum_{i,j} \log \frac{d^2(\mathbf{x}_i, \mathbf{y}_j)}{\epsilon^2} - \frac{c}{6} \sum_{i < j} \log \frac{d^2(\mathbf{x}_i, \mathbf{x}_j)}{\epsilon^2} - \quad (9)$$

$$- \frac{c}{6} \sum_{i < j} \log \frac{d^2(\mathbf{y}_i, \mathbf{y}_j)}{\epsilon^2}. \quad (10)$$

Main Results



Main Results

- ▶ Cauchy surfaces Σ are **finite-sized**, hence, do not need IR regularization.
- ▶ $S_m(\Sigma) \geq \frac{c}{3} \log \frac{2}{\kappa_c \varepsilon}$. **This result is in contradiction with pure state condition: $S_m(\Sigma) = 0$.**
- ▶ Entropies for a finite interval and its complement are

$$S_m(R) = \frac{c}{6} \log \left(\frac{2\sqrt{f(b)}}{\kappa_c^2 \varepsilon^2} [\cosh \kappa_c r_*(b) - \cosh \kappa_c (t_b - t_{0+})] \right)$$
$$S_m(\bar{R}) = \frac{c}{6} \log \left(\frac{2\sqrt{f(b)}}{\kappa_c^2 \varepsilon^2} [\cosh \kappa_c r_*(b) + \cosh \kappa_c (t_b - t_{0-})] \right)$$

Therefore, **complementarity is explicitly violated:**
 $S_m(R) \neq S_m(\bar{R})$.

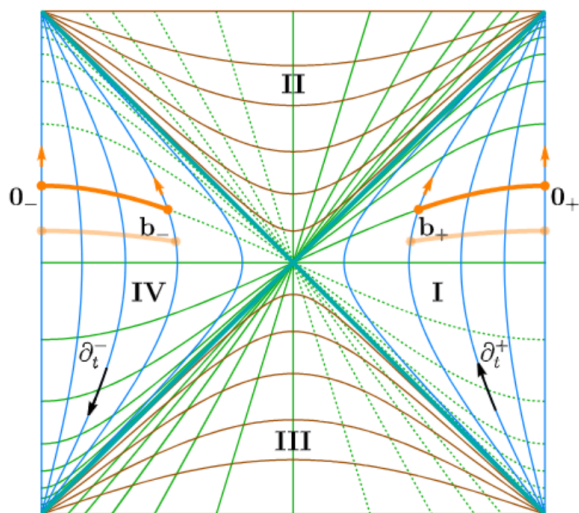
Main Results

Consider the finite region $R = [\mathbf{b}_+, \mathbf{0}_+]$.

- ▶ If $\kappa_c(t_b - t_{0_+}) = \text{const}$, then $S(R) = \text{const}$, and $S(R) \leq S_{\text{GH}}$.
- ▶ Let $\kappa_c(t_b - t_{0_+}) \equiv g(t_b) \neq \text{const}$. Then
 - ▶ If $g'(t_b) > 0$, then Cauchy surface breaks, $S(R) \rightarrow \infty$, and the problem becomes ill-defined.
 - ▶ If $g'(t_b) < 0$, the endpoints \mathbf{b}_+ and $\mathbf{0}_+$ asymptotically approach the hypersurface of constant time at late times, and $S(R) \rightarrow \text{const}$.

Main Results

Consider $R = [0_-, \mathbf{b}_-] \cup [\mathbf{b}_+, 0_+]$ with $t_{b_+} = -t_{b_-} \equiv t_b$,
 $t_{0_+} = t_{b_+}$, $t_{0_-} = t_{b_-}$.



Main Results

- ▶ At early times $t_b \ll r_*(b)/2$, the entropy monotonically increases

$$S_m(R) \simeq \frac{2c}{3} \kappa_c t_b + \text{const}$$

- ▶ At late times $t_b \gg r_*(b)/2$, the entropy saturates at the value

$$S_m(R) \simeq \frac{c}{3} \log \left(\frac{2\sqrt{f(b)}}{\kappa_c^2 \varepsilon^2} \right) + \frac{c}{3} \log (\cosh \kappa_c r_*(b) - 1)$$

- ▶ $\lim_{b \rightarrow \ell} r_*(b) \rightarrow \infty \implies$ as $b \rightarrow \ell$, the linear growth regime gets longer. **This might formally lead to the information paradox in dS.**
- ▶ Numerical analysis reveals **no island** for this configuration.

Conclusions

- ▶ As a consistency test of the setup, we studied basic properties of entanglement entropy for a pure state in dS. Both **pure state condition** and **complementarity** are **violated**. This disproves CFT₂ formulas for entanglement entropy to partially reduced dS geometry.
- ▶ We have given an example in this setup, which might **potentially lead to the information paradox**. Numerical calculations reveal **no non-trivial island** for this configuration. The **possible information paradox cannot be resolved** by the island formula.

Thank you for your attention!