

# Massive particle production in the field of massless plane wave

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- **Particle creation by a dynamic classical field. Example: the production of matter particles by the inflaton field oscillation at the end of inflation (reheating stage) [1]**
- **Particle production enhancement effect - parametric resonance**
- **Creation of massive particles by an intense wave of a massless field with resonant amplification [2], small mass  $m \ll \omega$**
- **The purpose of the work is to consider the case of arbitrary masses**

[1] Rubakov V.A. Gorbunov D.S. Introduction to the theory of the early universe. Cosmological perturbations. Inflationary theory. URSS, 2009.

[2] A.Arza. Phys.Rev.D 105 (2022) arXiv:2009.03870

# $\mathcal{H} = g\phi\chi^2$ Two-scalar model

## Equations of motion

$$\square\phi = -g\chi^2,$$

Monochromatic plane wave

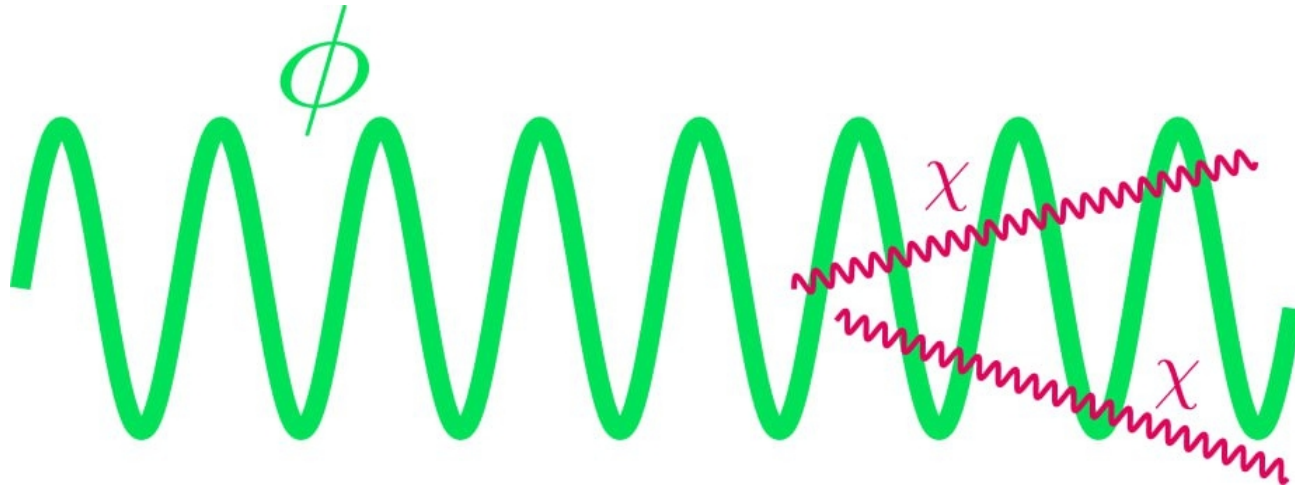
$$(\square + m_\chi^2)\chi = -2g\phi\chi \quad \phi(\vec{x}, t) = \frac{\sqrt{2\rho_\phi}}{\omega_p} \cos(\vec{p} \cdot \vec{x} - \omega_p t)$$

$$\Omega_{\vec{k}} = \sqrt{k^2 + m_\chi^2} \quad \omega_{\vec{p}} = \sqrt{p^2 + m_\phi^2}$$

Quantum field

$$\chi = \int \frac{d^3k}{(2\pi)^3} \frac{1}{\sqrt{2\Omega_{\vec{k}}}} \left( \chi_{\vec{k}}(t) e^{i\vec{k} \cdot \vec{x}} + \chi_{\vec{k}}(t)^\dagger e^{-i\vec{k} \cdot \vec{x}} \right)$$

$$\mathcal{H} = g\phi\chi^2$$



**Denote**

$$A_{\vec{k}} = \chi_{\vec{k}} + \chi_{-\vec{k}}^\dagger$$

$$(\partial_t^2 + \Omega_{\vec{k}}^2)A_{\vec{k}} = -\omega_{\vec{p}}^2\alpha \left( \sqrt{\frac{\Omega_{\vec{k}}}{\Omega_{\vec{k}-\vec{p}}}}A_{\vec{k}-\vec{p}}e^{-i\omega_{\vec{p}}t} + \sqrt{\frac{\Omega_{\vec{k}}}{\Omega_{\vec{k}+\vec{p}}}}A_{\vec{k}+\vec{p}}e^{i\omega_{\vec{p}}t} \right)$$

where  $\alpha \equiv \frac{g\sqrt{2\rho_\phi}}{\omega_{\vec{p}}^3}$

**Suppose**  $\chi_{\vec{k}} = a_{\vec{k}}(t)e^{-i\Omega_{\vec{k}}t}$

$$\begin{aligned}
& e^{-i\Omega_{\vec{k}}t}(\ddot{a}_{\vec{k}} - 2i\Omega_{\vec{k}}\dot{a}_{\vec{k}}) + e^{i\Omega_{-\vec{k}}t}(\ddot{a}_{-\vec{k}}^\dagger + 2i\Omega_{-\vec{k}}\dot{a}_{-\vec{k}}^\dagger) = \\
& = -\omega_{\vec{p}}^2\alpha \left( \sqrt{\frac{\Omega_{\vec{k}}}{\Omega_{\vec{k}+\vec{p}}}} \left( a_{-\vec{k}-\vec{p}}^\dagger e^{i(\Omega_{-\vec{k}-\vec{p}}+\omega_{\vec{p}})t} + a_{\vec{k}+\vec{p}} e^{-i(\Omega_{\vec{k}+\vec{p}}-\omega_{\vec{p}})t} \right) + \right. \\
& \quad \left. + \sqrt{\frac{\Omega_{\vec{k}}}{\Omega_{\vec{k}-\vec{p}}}} \left( a_{-\vec{k}+\vec{p}}^\dagger e^{i(\Omega_{-\vec{k}+\vec{p}}-\omega_{\vec{p}})t} + a_{\vec{k}-\vec{p}} e^{-i(\Omega_{\vec{k}-\vec{p}}+\omega_{\vec{p}})t} \right) \right)
\end{aligned}$$

The amplitude  $a_{-\vec{k}+\vec{p}}^\dagger$  - leading,  $a_{-\vec{k}-\vec{p}}^\dagger$  - subleading

# The final equation

- **Approximation (A.Arza, PRD 2022):  $\ddot{a}$  term neglected  $\rightarrow$**

$$m_\chi \ll \omega_p$$

$$\dot{a}_{\vec{k}} = -i\sigma_{\vec{k}} a_{\vec{p}-\vec{k}}^\dagger e^{i\epsilon_{\vec{k}} t}$$

$$\sigma_{\vec{k}} = g \sqrt{\frac{\rho_\phi/2}{\omega_{\vec{p}}^2 \Omega_{\vec{p}-\vec{k}} \Omega_{\vec{k}}}}$$

- **Without approximation**

$$e^{-i\Omega_{\vec{k}} t} (\ddot{a}_{\vec{k}} - 2i\Omega_{\vec{k}} \dot{a}_{\vec{k}}) = \sigma_{p-k} a_{-\vec{k}+\vec{p}}^\dagger e^{i(\Omega_{-\vec{k}+\vec{p}} - \omega_{\vec{p}}) t}$$

$$\sigma_{p-k} = -\omega_p^2 \alpha \sqrt{\frac{\Omega_{\vec{k}}}{\Omega_{\vec{k}-\vec{p}}}}$$

# Solution

- In the approximation**

$$\epsilon_{p-k} = \Omega_k + \Omega_{p-k} - \omega_p = \epsilon_k$$

$$a_{\vec{k}}(t) = e^{i\epsilon_{\vec{k}}t/2} \left( a_{\vec{k}}(0) \left( \cosh(s_{\vec{k}}^0 t) - i \frac{\epsilon_{\vec{k}}}{2s_{\vec{k}}^0} \sinh(s_{\vec{k}}^0 t) \right) - i \frac{\sigma_{\vec{p}-\vec{k}}}{2s_{\vec{k}}^0 \Omega_{\vec{k}}} a_{\vec{p}-\vec{k}}^\dagger(0) \sinh(s_{\vec{k}}^0 t) \right)$$

- Without approximation**

$$s_{\vec{k}}^0 = \frac{1}{2} \sqrt{\frac{\sigma_{\vec{p}-\vec{k}}^2}{\Omega_{\vec{k}}^2} - \epsilon_{\vec{k}}^2}$$

$$a_{\vec{k}}(t) = e^{i\epsilon_{\vec{k}}t/2} \left[ a_{\vec{k}}(0) \left( \cosh(s_{\vec{k}} t) - i \frac{\epsilon_{\vec{k}}^2/4 - s_{\vec{k}}^2 - \Omega_{\vec{k}} \epsilon_{\vec{k}}}{s_{\vec{k}}(2\Omega_{\vec{k}} - \epsilon_{\vec{k}})} \sinh(s_{\vec{k}} t) \right) - a_{\vec{p}-\vec{k}}^\dagger(0) \cdot i \frac{\sigma_{\vec{p}-\vec{k}}}{s_{\vec{k}}(2\Omega_{\vec{k}} - \epsilon_{\vec{k}})} \sinh(s_{\vec{k}} t) \right],$$

$$s_{\vec{k}}^2 = \sqrt{\Omega_{\vec{k}}^2 (2\Omega_{\vec{k}} - \epsilon_{\vec{k}})^2 + \sigma_{\vec{p}-\vec{k}}^2} - \Omega_{\vec{k}} (2\Omega_{\vec{k}} - \epsilon_{\vec{k}}) - \frac{\epsilon_{\vec{k}}^2}{4}$$

# Production of $\chi$ modes in the external $\varphi$ field

$$s^2 > 0$$

## The occupancy number

$$f_{\chi, \vec{k}}(t) = \langle 0 | a_{\vec{k}}^\dagger(t) a_{\vec{k}}(t) | 0 \rangle = \frac{\sigma_{\vec{p}-\vec{k}}^2}{4} \times \begin{cases} \frac{\sinh^2(s_{\vec{k}}^0 t)}{\left(s_{\vec{k}}^0\right)^2 \Omega_{\vec{k}}^2}, \\ \frac{\sinh^2(s_{\vec{k}} t)}{s_{\vec{k}}^2 (\Omega_{\vec{k}} - \epsilon_{\vec{k}}/2)^2} \end{cases}$$

## The total density

$$n_\chi(t) = \int \frac{d^3 k}{(2\pi)^3} f_{\chi, \vec{k}}(t)$$



# Dimensionless parameters

$$\eta_{\vec{k}} = s_{\vec{k}} / \omega_{\vec{p}}$$

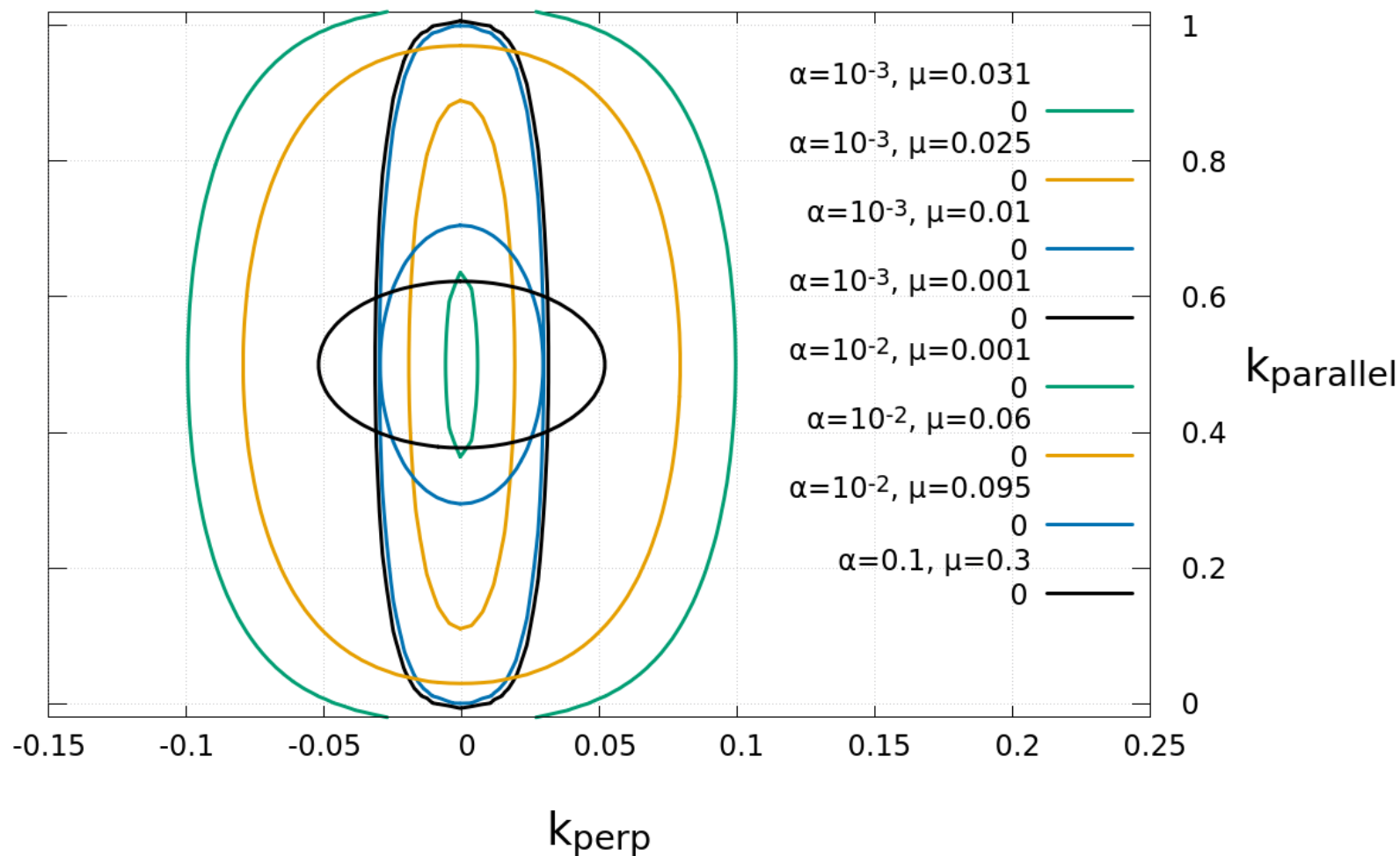
$$\vec{k} = \vec{k}_{\parallel} + \vec{k}_{\perp} \quad \frac{\Omega_{\vec{k}}}{\omega_{\vec{p}}} \equiv \beta_{\vec{k}} = \sqrt{\kappa_{\parallel}^2 + \kappa_{\perp}^2 + \mu^2}$$

$$\frac{\Omega_{\vec{p}-\vec{k}}}{\omega_{\vec{p}}} \equiv \beta_{\vec{v}-\vec{k}} = \sqrt{1 - 2\kappa_{\parallel} + \kappa_{\parallel}^2 + \kappa_{\perp}^2 + \mu^2}$$

$$\eta_{\vec{k}}^2 = \begin{cases} \frac{\alpha^2}{4\beta_{\vec{k}}\beta_{\vec{v}-\vec{k}}} - \frac{(\beta_{\vec{k}} + \beta_{\vec{v}-\vec{k}} - 1)^2}{4}, & \text{(approximation)} \\ \sqrt{\beta_{\vec{k}}^2 (\beta_{\vec{k}} - \beta_{\vec{v}-\vec{k}} + 1)^2 + \alpha^2 \beta_{\vec{k}} \beta_{\vec{v}-\vec{k}}^{-1}} - \beta_{\vec{k}} (\beta_{\vec{k}} - \beta_{\vec{v}-\vec{k}} + 1) - \frac{(\beta_{\vec{k}} + \beta_{\vec{v}-\vec{k}} - 1)^2}{4} \end{cases}$$

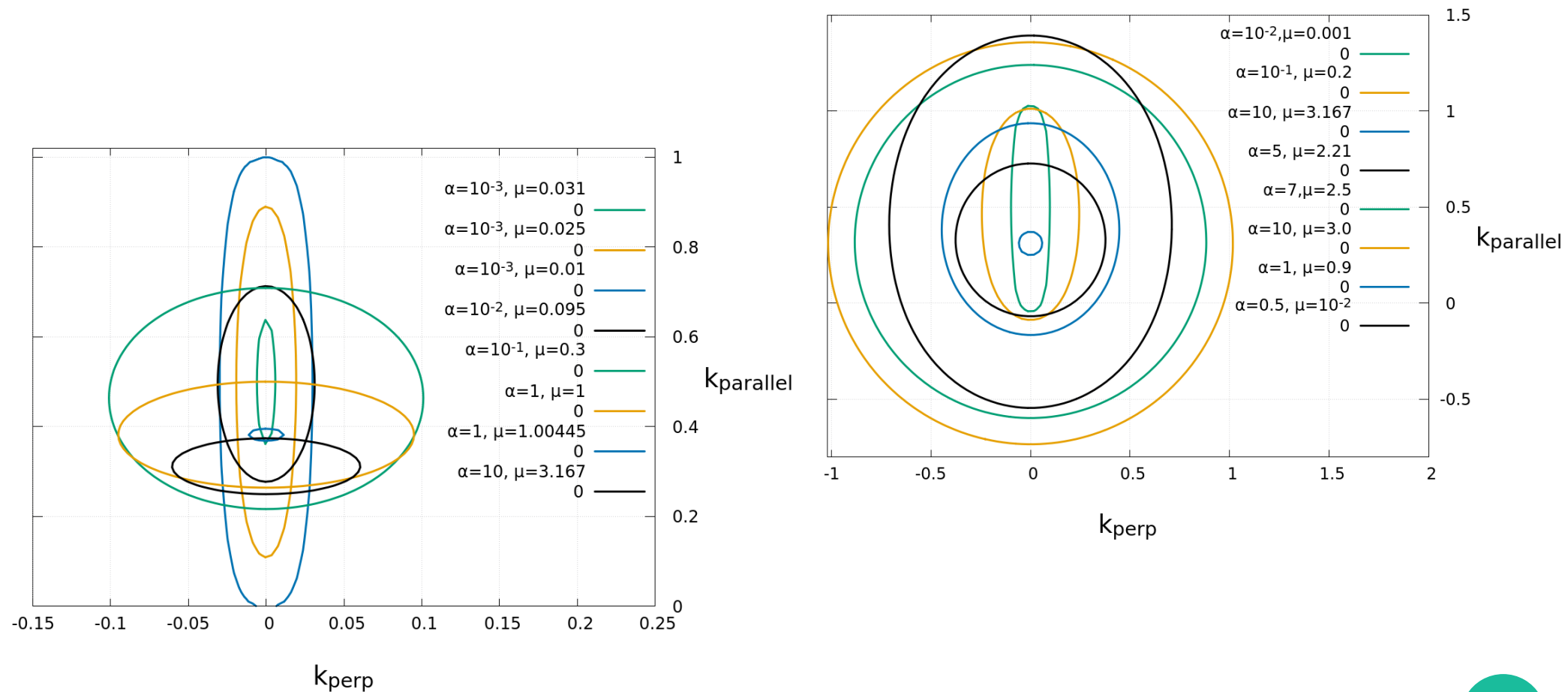
# The contours of $\eta_k = 0$ as the function of parallel and perpendicular momenta

## Approximation case



# The contours of $\eta_k = 0$ as the function of parallel and perpendicular momenta

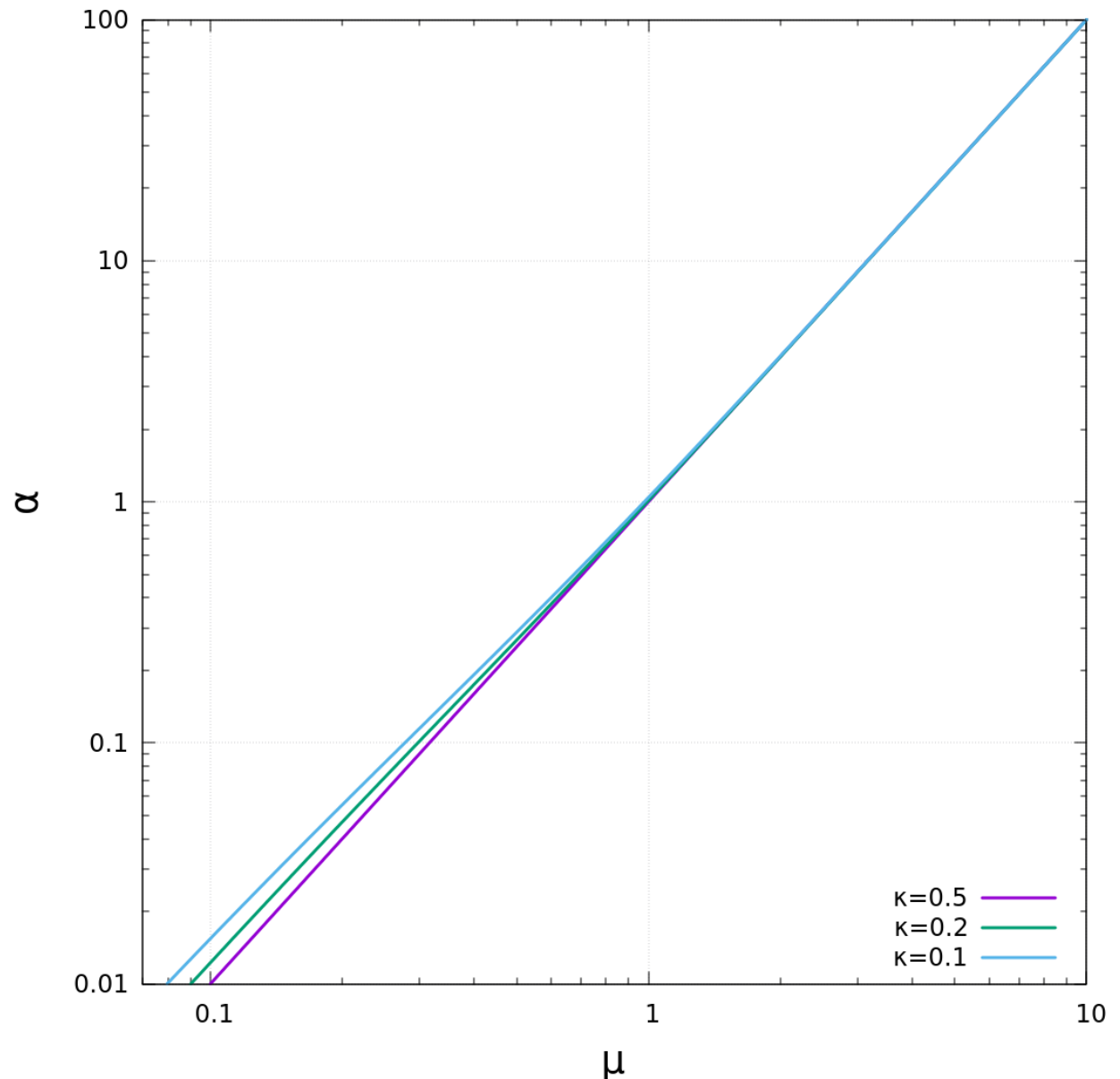
## Without approximation



# The boundary of stability

$$\eta_{\vec{k}} = 0$$

$$\rho_{\phi} \geq \frac{m_{\chi}^4 \omega_p^2}{2g^2}$$



$$\alpha(\mu) = 4 \sqrt{\frac{\beta_{v-k}}{\beta_k} (\beta_k + \beta_{v-k} - 1)^2 (-3\beta_k + \beta_{v-k} - 1)^2}$$

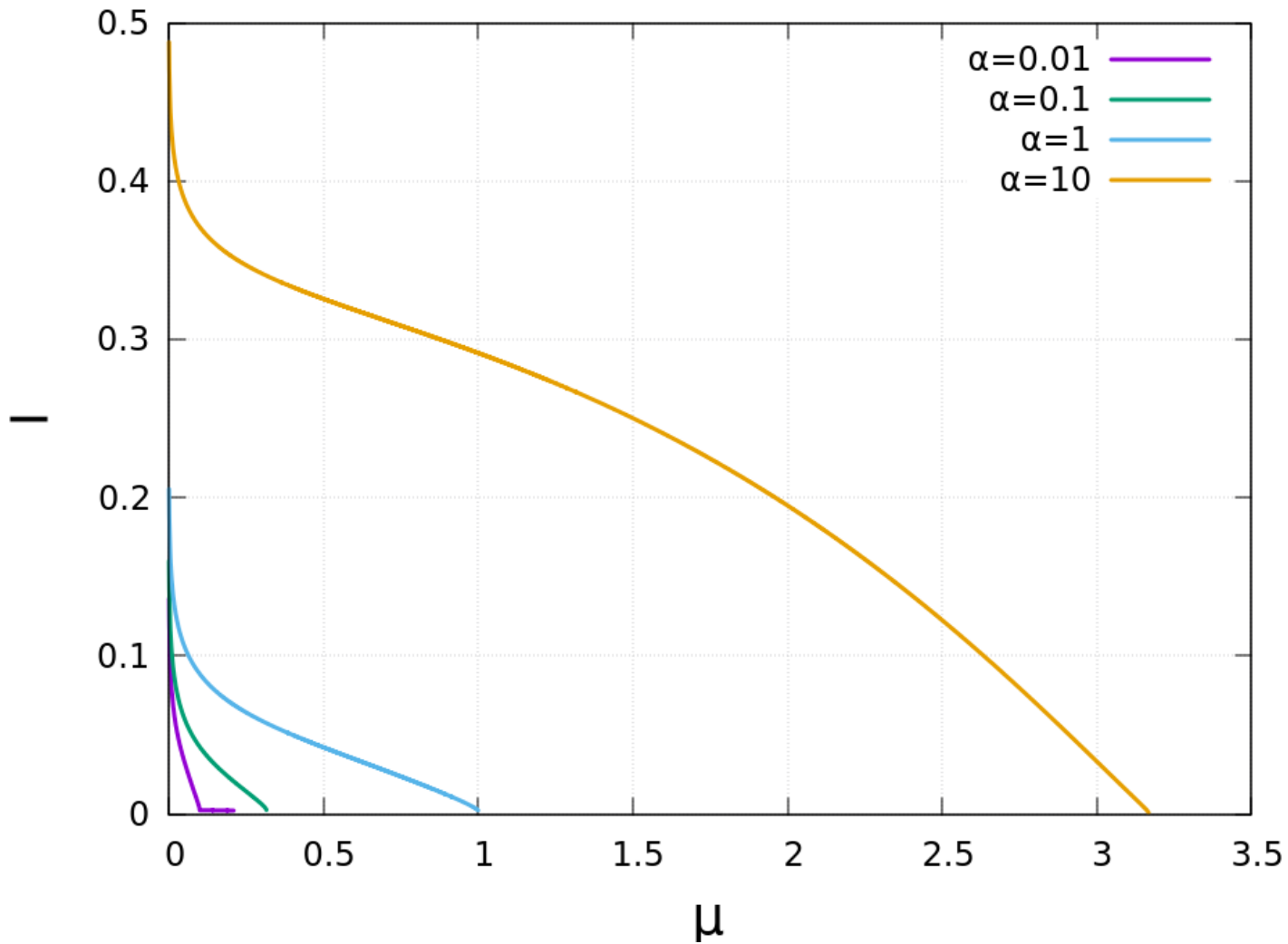
# Density of modes

$$n_{\chi}(t) = \int \frac{d^3 k}{(2\pi)^3} f_{\chi, \vec{k}}(t)$$

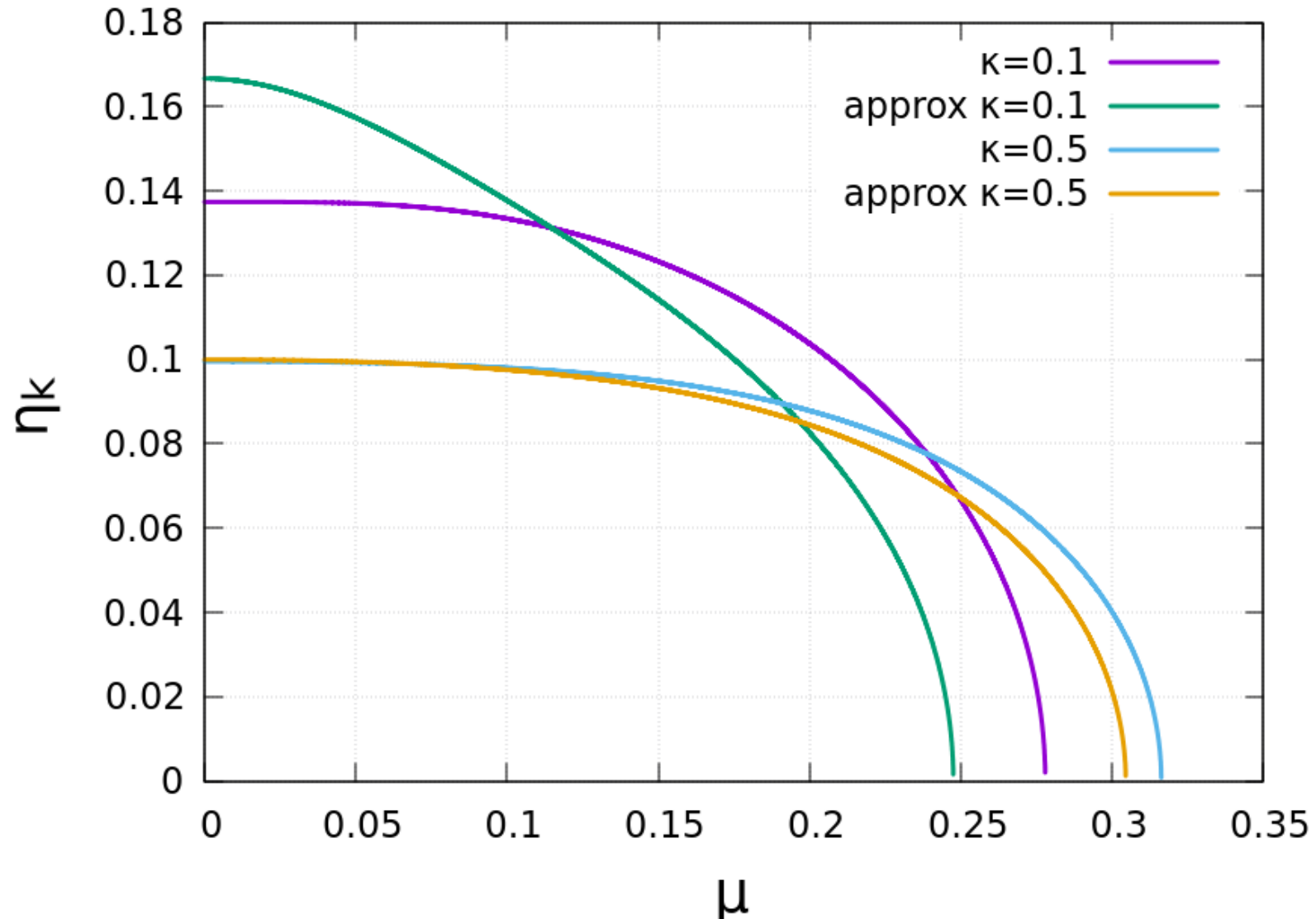
$$n_{\chi}(t) = \frac{\alpha^2 t^2 \omega_p^5}{(2\pi)^2} \int \kappa_{\perp} d\kappa_{\perp} I(\alpha, \mu, \kappa_{\perp})$$

$$n_{\chi}(t) = \alpha^2 t^2 \omega_p^5 \frac{\int d\kappa_{\parallel} \int \kappa_{\perp} d\kappa_{\perp}}{(2\pi)^2} \times \begin{cases} \frac{1}{4\beta_{\vec{k}} \beta_{\vec{v}-\vec{k}}}, & \text{(approximation)} \\ \frac{\beta_{\vec{k}}}{\beta_{\vec{v}-\vec{k}} (\beta_{\vec{k}} - \beta_{\vec{v}-\vec{k}} + 1)^2}. & \text{(no approx)} \end{cases}$$

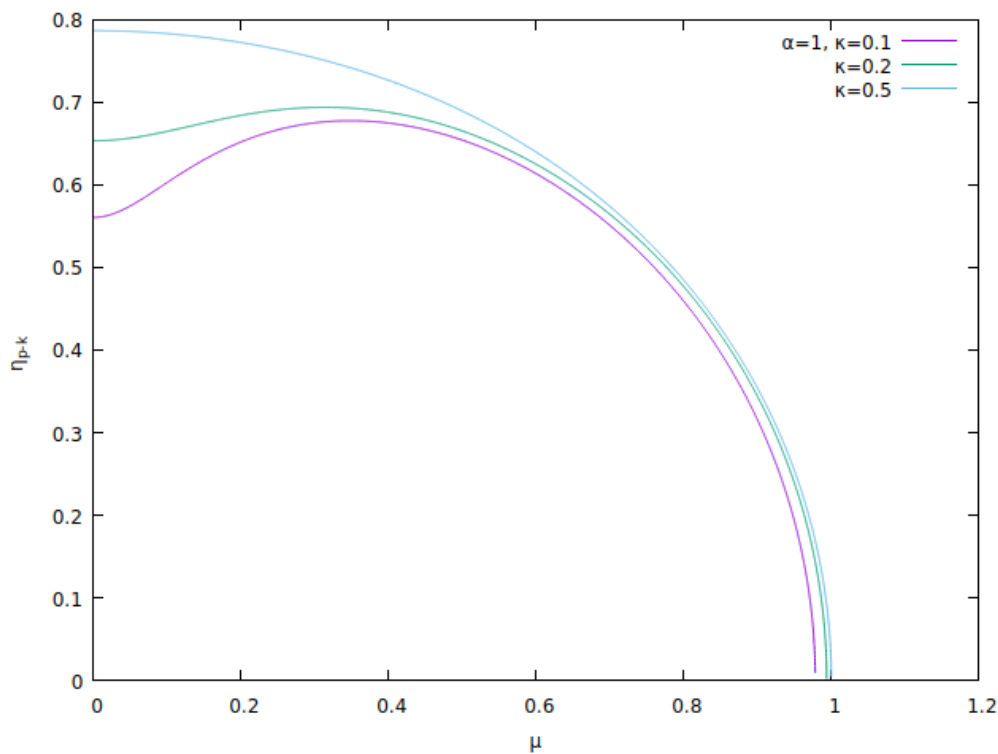
# Dependence $I(\mu)$ for fixed $\alpha$



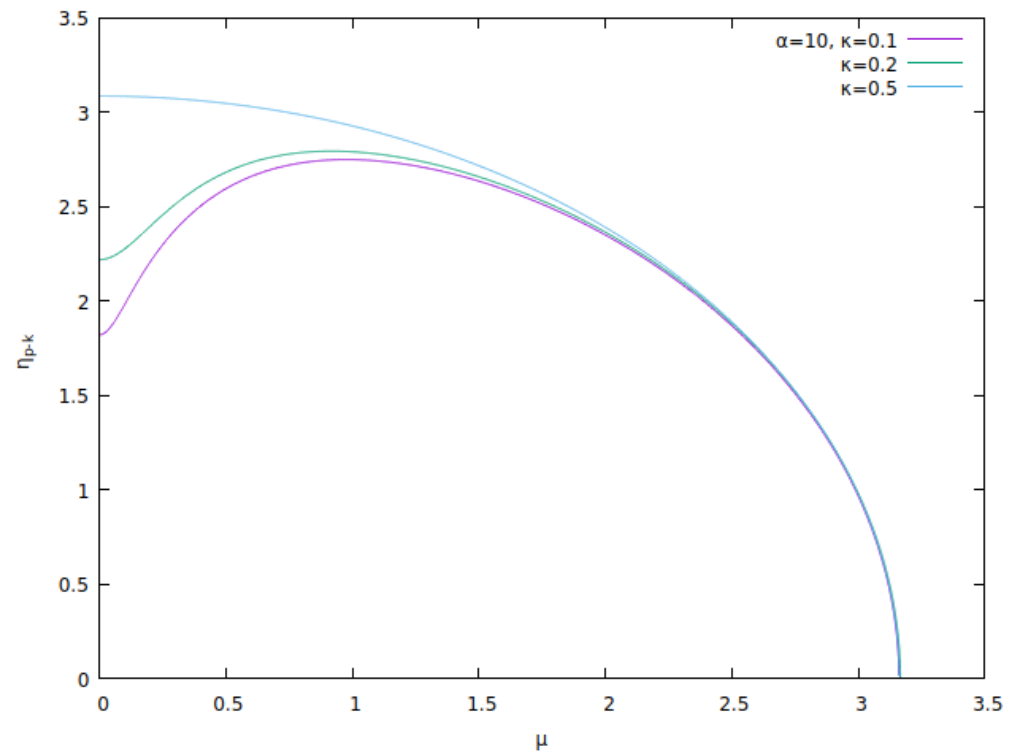
# Dependence of $\eta_k$ on $\mu$ at $\alpha=0.1$ and $\kappa=0.1, 0.5$ in the approximation and without it



# Dependences of $\eta_k$ on $\mu$ for various $\alpha$ (without approximation)



$\alpha=1$



$\alpha=10$



Solution for  $a_{-\vec{k}}^\dagger$

$$e^{i\Omega_{-\vec{k}}t} (\ddot{a}_{-\vec{k}}^\dagger + 2i\Omega_{-\vec{k}}\dot{a}_{-\vec{k}}^\dagger) = \sigma_{\vec{k}+\vec{p}} a_{\vec{k}+\vec{p}} e^{-i(\Omega_{\vec{k}+\vec{p}} - \omega_{\vec{p}})t}$$

$$a_{-\vec{k}}^\dagger = e^{-i\epsilon_{-\vec{k}}t/2} \left[ a_{-\vec{k}}^\dagger(0) (\cosh(s_{-\vec{k}}t) + i \frac{-\frac{\epsilon_{-\vec{k}}^2}{4} + s_{-\vec{k}}^2 + \Omega_{\vec{k}}\epsilon_{-\vec{k}}}{s_{-\vec{k}}(-\epsilon_{-\vec{k}} + 2\Omega_{\vec{k}})} \sinh(s_{-\vec{k}}t)) + i \frac{-\sigma_{\vec{p}+\vec{k}}}{s_{-\vec{k}}(-\epsilon_{-\vec{k}} + 2\Omega_{\vec{k}})} a_{\vec{p}+\vec{k}}(0) \sinh(s_{-\vec{k}}t) \right]$$

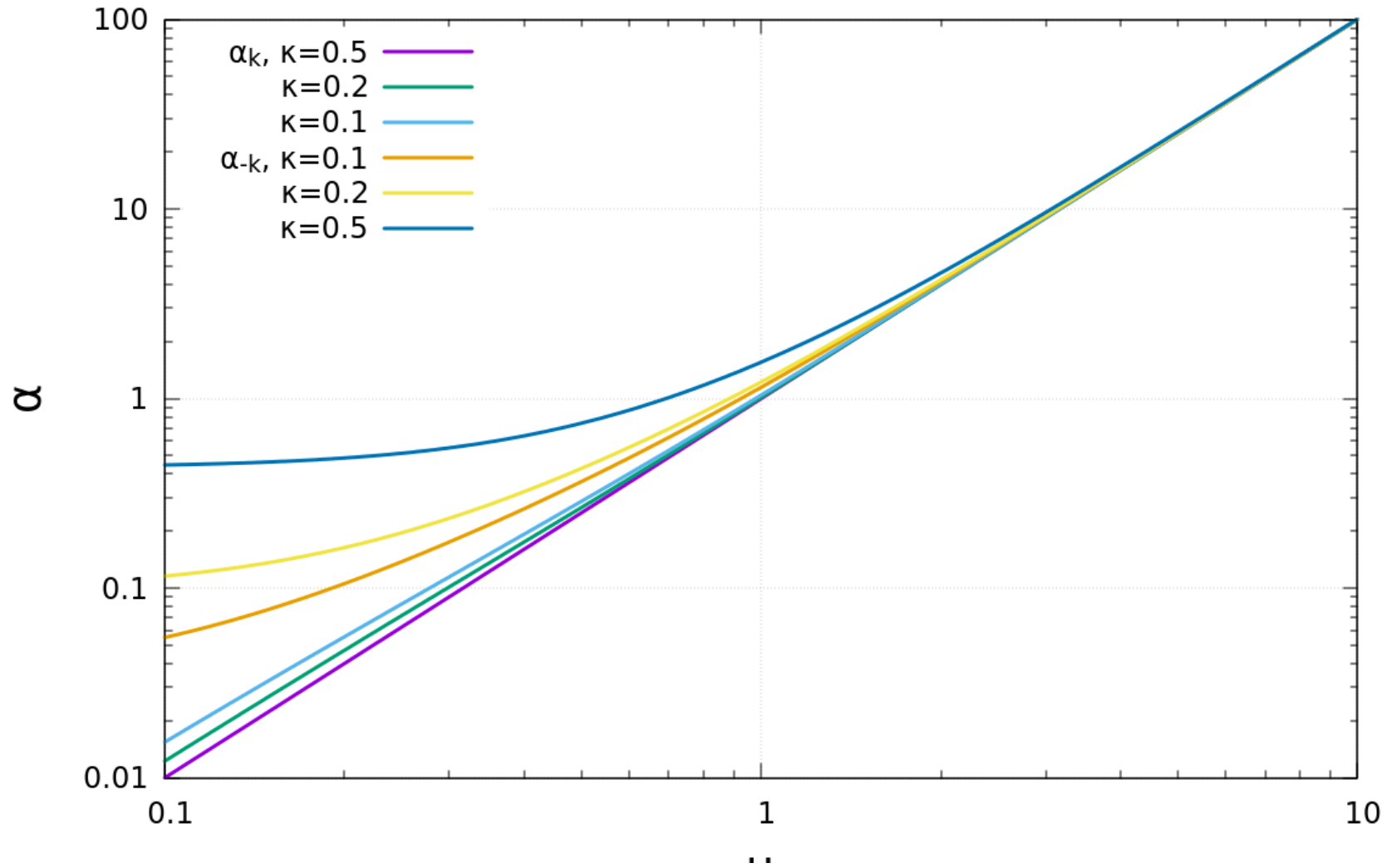
$$\sigma_{\vec{k}+\vec{p}} = -\omega_{\vec{p}}^2 \alpha \sqrt{\frac{\Omega_{\vec{k}}}{\Omega_{\vec{k}+\vec{p}}}}$$

$$\epsilon_{-\vec{k}} = \Omega_{\vec{k}} + \Omega_{\vec{p}+\vec{k}} - \omega_{\vec{p}}$$

$$s_{-\vec{k}}^2 = -\frac{\epsilon_{-\vec{k}}^2}{4} - 2\Omega_{\vec{k}}^2 + \epsilon_{-\vec{k}}\Omega_{\vec{k}} + \sqrt{\Omega_{\vec{k}}^2\epsilon_{-\vec{k}}^2 + 4\Omega_{\vec{k}}^4 + \sigma_{\vec{p}+\vec{k}}^2 - 4\epsilon_{-\vec{k}}\Omega_{\vec{k}}^3}$$

# The boundary of stability for $a_k$ and

$a_{-k}^+$



$$\alpha_{-\kappa}(\mu) = \frac{1}{4} \sqrt{\frac{\beta_{\vec{v}+\vec{\kappa}}}{\beta_{\vec{\kappa}}} (\beta_{\vec{\kappa}} + \beta_{\vec{v}+\vec{\kappa}} - 1)^2 (-3\beta_{\vec{\kappa}} + \beta_{\vec{v}+\vec{\kappa}} - 1)^2}$$

# Standing wave

$$\phi(\xi, t) = 8 \frac{\sqrt{2\rho\phi}}{\omega} \cos(p_x \xi) \cos(p_y \xi) \cos(p_z \xi) \cos(\omega t)$$

**The equation** 
$$e^{-i\Omega_{\vec{k}} t} (\ddot{a}_{\vec{k}} - 2i\Omega_{\vec{k}} \dot{a}_{\vec{k}}) = \sigma_{\vec{p}_j - \vec{k}} a_{\vec{p}_j - \vec{k}}^\dagger e^{i(\Omega_{\vec{p}_j - \vec{k}} - \omega)t}$$

$j = 1, \dots, 4$

## Solution

$$a_{\vec{k}}(t) = \sigma_{\vec{p}_j - \vec{k}} e^{i\epsilon_{\vec{p}_j - \vec{k}} t/2} \left[ a_{\vec{k}}(0) (\cosh(s_{\vec{p}_j - \vec{k}} t) - i \frac{\frac{p_j^2}{4} - s_{\vec{p}_j - \vec{k}}^2 - \Omega_{\vec{k}} \epsilon_{\vec{p}_j - \vec{k}}}{s_{\vec{p}_j - \vec{k}} (\epsilon_{\vec{p}_j - \vec{k}} - 2\Omega_{\vec{k}})} \sinh(s_{\vec{p}_j - \vec{k}} t)) - i \frac{\sigma_{\vec{p}_j - \vec{k}}}{s_{\vec{p}_j - \vec{k}} (\epsilon_{\vec{p}_j - \vec{k}} - 2\Omega_{\vec{k}})} a_{\vec{p}_j - \vec{k}}^\dagger(0) \sinh(s_{\vec{p}_j - \vec{k}} t) \right]$$

$$\vec{\xi} = (\xi_x, \xi_y, \xi_z), \vec{p}_1 = (p_x, p_y, p_z), \vec{p}_2 = (p_x, -p_y, p_z)$$

$$\vec{p}_3 = (-p_x, -p_y, p_z), \vec{p}_4 = (-p_x, p_y, p_z)$$

# Conclusions

- **The decay of  $\varphi \rightarrow \chi\chi$  occurs at the amplitude of the field  $\varphi$  above the threshold, not only in the case of small masses ( $m_\chi \ll \omega_p$ ) - the case of A.Arza, but also at an arbitrary mass**
- **In the case of large masses ( $m_\chi \gg \omega_p$ ), the required threshold amplitude of the field  $\varphi$  is powerfully greater in comparison with the case of small masses**

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**Thanks for your  
attention!**