Antistars in Milky Way

A. D. Dolgov

Novosibirsk State University, Novosibirsk, Russia

Internatonal Seminar Quarks 2020



- 1. Announcement of possible antistar registration.
- 2. Antistar dark matter.
- 3.

Possible discovery of several **antistars in the Galaxy** was recently reported: "Constraints on the antistar fraction in the Solar System neighborhood from the 10-year Fermi Large Area Telescope gamma-ray source catalog." S. Dupourqué, L. Tibaldo, P. Von Ballmoos, Phys.Rev.D 103 (2021) 8, 083016 • e-Print: 2103.10073 [astro-ph.HE] "We identify in the catalog 14 antistar candidates not associated with any objects belonging to established gamma-ray source classes and with a spectrum compatible with bar:yon-antibaryon annihilation."

Antistar prediction

Existence of antistars in the Galaxy was predicted many years ago. Mechanism of massive PBH formation with log-normal mass spectrum: A. Dolgov and J.Silk, PRD 47 (1993) 4244 "Baryon isocurvature fluctuations at small scale and baryonic dark matter. A.Dolgov, M. Kawasaki, N. Kevlishvili, Nucl. Phys. B807 (2009) 229, "Inhomogeneous baryogenesis, cosmic antimatter, and dark matter". Massive PBHs allow to cure multpie inconsistencies with the standard cosmology and astrophysics. Unusual stellar type compact objects are also created, including abundant antistars in the Galaxy. The chirp mass distribution of LIGO events very well agrees with theoretical predictions, AD, A.G. Kuranov, N.A. Mitichkin, S. Porey, K.A. Postnov, et al JCAP 12 (2020) 017 - a strong support of the scenario.

Bounds on antistars in the Galaxy

As argued in:

C. Bambi, A.D. Dolgov, "Antimatter in the Milky Way", Nucl.Phys.B 784 (2007) 132-150 • e-Print: astro-ph/0702350,

A.D. Dolgov, S.I. Blinnikov, "Stars and Black Holes from the very Early Universe", Phys.Rev.D 89 (2014) 2, 021301 • e-Print: 1309.3395, S.I. Blinnikov, A.D. Dolgov, K.A. Postnov, "Antimatter and antistars in the universe and in the Galaxy", Phys.Rev.D 92 (2015) 2, 023516 • e-print: 1409.5736

an abundant density of compact anti-stars in the universe and even in the Galaxy does not violate existing observational limits.

In a recent publication a striking idea was put forward that dark matter may consist of compact anti-stars: J. S. Sidhu, R.J. Scherrer, G. Starkman, "Antimatter as Macroscopic Dark Matter", arXiv:2006.01200, astro-ph.CO. Such anti-DM may be easier to spot than other forms of macroscopic DM.

Antimatter creation by mirror matter

Antistars or antimatter cores in mirror neutron stars? Zurab Berezhiani (Jun 21, 2021) e-Print: 2106.11203 [astro-ph.HE]

The oscillation of the neutron n into mirror neutron n', its partner from dark mirror sector, can gradually transform an ordinary neutron star into a mixed star consisting in part of mirror dark matter. The implications of the reverse process taking place in the mirror neutron stars depend on the sign of baryon asymmetry in mirror sector. Namely, if it is negative, as predicted by certain baryogenesis scenarios, then $\bar{n}' - \bar{n}$ transitions create a core of our antimatter gravitationally trapped in the mirror star interior.

Old Anti-Creation Mechanism

The mechanism of massive PBH and antistar formation: A. Dolgov and J.Silk, PRD 47 (1993) 4244 "Baryon isocurvature fluctuations at small scale and baryonic dark matter. A.Dolgov, M. Kawasaki, N. Kevlishvili, Nucl.Phys. B807 (2009) 229, "Inhomogeneous baryogenesis, **cosmic antimatter**, and dark matter". Massive PBHs allow to cure multpie inconsistencies with the standard cosmology and astrophysics.

Unusual stellar type compact objects could also be created. The model predicts the log-normal mass spectrum of PBH:

$$\frac{dN}{dM} = \mu^2 \exp\left[-\gamma \ln^2(M/M_0)\right],$$

and predicts $M_0 \approx 10 M_{\odot}$. A.Dolgov, K.Postnov, Why the mean mass of primordial black hole distribution is close to $10 M_{\odot}$. JCAP 07 (2020) 063 • e-Print: 2004.11669 . Very well agrees with observations.

SUSY motivated baryogenesis, Affleck and Dine (AD).

SUSY predicts existence of scalars with $B \neq 0$. Such bosons may condense along flat directions of the quartic potential:

 $U_{\lambda}(\chi) = \lambda |\chi|^4 \left(1 - \cos 4\theta\right)$

and of the mass term, $U_m = m^2 \chi^2 + m^{*\,2} \chi^{*\,2}$:

$$U_m(\chi) = m^2 |\chi|^2 [1 - \cos\left(2\theta + 2\alpha\right)],$$

where $\chi = |\chi| \exp(i\theta)$ and $m = |m|e^{\alpha}$. If $\alpha \neq 0$, C and CP are broken. In GUT SUSY baryonic number is naturally non-conserved - non-invariance of $U(\chi)$ w.r.t. phase rotation.

Anti-Creation Mechanism

Initially (after inflation) χ is away from origin and, when inflation is over, starts to evolve down to equilibrium point, $\chi = 0$, according to Newtonian mechanics:

$$\ddot{\chi} + \mathbf{3H}\dot{\chi} + \mathbf{U}'(\chi) = \mathbf{0}.$$

Baryonic charge of χ :

 $B_{\chi} = \dot{\theta} |\chi|^2$

is analogous to mechanical angular momentum. χ decays transferred baryonic charge to that of quarks in B-conserving process.

AD baryogenesis could lead to baryon asymmetry of order of unity, much larger than the observed 10^{-9} .

Anti-Creation Mechanism

If $m \neq 0$, the angular momentum, B, is generated by a different direction of the quartic and quadratic valleys at low χ . If CP-odd phase α is small but non-vanishing, both baryonic and antibaryonic domains might be formed with possible dominance of one of them. Matter and antimatter domains may exist but globally $B \neq 0$.

Affleck-Dine field χ with CW potential coupled to inflaton Φ (AD and Silk; AD, Kawasaki, Kevlishvili):

$$U = g|\chi|^2 (\Phi - \Phi_1)^2 + \lambda|\chi|^4 \ln\left(\frac{|\chi|^2}{\sigma^2}\right)$$
$$+\lambda_1(\chi^4 + h.c.) + (m^2\chi^2 + h.c.).$$

Coupling to inflaton is the general renormalizable one. When the window to the flat direction is open, near $\Phi = \Phi_1$, the field χ slowly diffuses to large value, according to quantum diffusion equation derived by Starobinsky, generalized to a complex field χ .

A. D. Dolgov

Anti-Creation Mechanism

If the window to flat direction, when $\Phi \approx \Phi_1$ is open only during a short period, cosmologically small but possibly astronomically large bubbles with high β could be created, occupying a small fraction of the universe, while the rest of the universe has normal $\beta \approx 6 \cdot 10^{-10}$, created by small χ . The mechanism of massive PBH formation quite different from all others. The fundament of PBH creation is build at inflation by making large isocurvature fluctuations at relatively small scales, with practically vanishing density perturbations.

Initial isocurvature perturbations are in chemical content of massless quarks. Density perturbations are generated rather late after the QCD phase transition.

The emerging universe looks like a piece of Swiss cheese, where holes are high baryonic density objects occupying a minor fraction of the universe volume.

Evolution of AD-field potential

Effective potential of χ for different values of the inflaton field Φ . The upper blue curve corresponds to a large value $\Phi > \Phi 1$ which gradually decreases down to $\Phi =$ $\Phi 1$, red curve. Then the potential returns back to the almost initial shape, as Φ drops down to zero. The evolution of χ in such a potential is similar to a motion of a pointlike particle (shown as a black ball in the figure) in Newtonian mechanics. First, due to quantum initial fluctuations χ left the unstable extremum of the potential at $\chi = 0$ and "tried" to keep pace with the moving potential minimum and later started to oscillate around it with decreasing amplitude. The decrease of the oscillation amplitude was induced by the cosmological expansion. In mechanical analogy the effect of the expansion is equivalent to the liquid friction term, $3H\chi'$. When Φ dropped below $\Phi 1$, the potential recovered its original form with the minimum at $\chi = 0$ and χ ultimately returned to zero but before that it could give rise to a large baryon asymmetry

$$\ddot{\chi} + 3H\dot{\chi} + U'(\chi) = 0.$$

Evolution of AD-field

(Dolgov - Kawasaki - Kevlishvili)

Field χ "rotates" in this plane with quite large angular momentum, which exactly corresponds to the baryonic number density of χ . Later χ decayed into quarks and other particles creating a large cosmological baryon asymmetry.



Results of (Anti-)Creation

The outcome, depending on $\beta = n_B/n_{\gamma}$.

- PBHs with log-normal mass spectrum.
- Compact stellar-like objects, as e.g. cores of red giants.
- Disperse hydrogen and helium clouds with (much) higher than average n_B density.
- β may be negative leading to creation of (compact?) antistars which could survive annihilation with the homogeneous baryonic background.

SUMMARY

- 1. Natural baryogenesis model leads to abundant fomation of PBHs and compact stellar-like objects in the early universe after QCD phase transition, $t \gtrsim 10^{-5}$ sec.
- 2. Predicted Log-normal mass spectrum of PBHs, confirmed by observations.
- 3. PBHs formed at this scenario can explain the peculiar features of the sources of GWs observed by LIGO.
- 4. The considered mechanism solves the numerous mysteries of $z \sim 10$ universe: abundant population of supermassive black holes, early created gamma-bursters and supernovae, early bright galaxies, and evolved chemistry including dust.
- 5. There is persuasive data in favor of the inverted picture of galaxy formation, when first a supermassive BH seeds are formed and later they accrete matter forming galaxies.

SUMMARY

- 6. An existence of supermassive black holes observed in all large and some small galaxies and even in almost empty environment is naturally explained.
- 7. "Older than *t_U*" stars may exist; the older age is mimicked by the unusual initial chemistry.
- 8. Existence of high density invisible "stars" (machos) is understood.
- 9. Explanation of origin of BHs with 2000 M_{\odot} in the core of globular cluster and the observed density of GCs is presented.
- 10. A large number of the recently observed IMBH was predicted.
- 11. A large fraction of dark matter or 100% can be made of PBHs.
- 12. Clouds of matter with high baryon-to-photon ratio.
- 13. A possible by-product: plenty of (compact) anti-stars, even in the Galaxy, not yet excluded by observations.

More data are badly needed

THE END or BEGINNING