

# No stable static spherically symmetric wormholes in Horndeski theory



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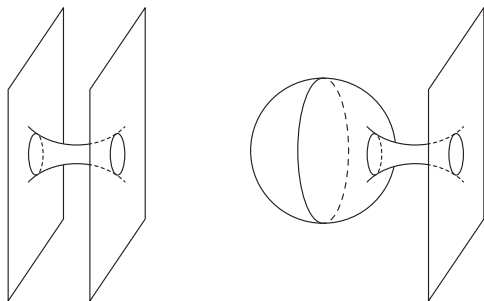
May 31, 2018



- 1 **Introduction**
  - Wormholes and NEC
- 2 **Stability conditions**
  - Odd sector
  - Even sector
- 3 **No-go theorem**
- 4 **Conclusions**



## Wormholes and semiclosed worlds



### Penrose theorem

- Noncompactness of the Cauchy hypersurface.
- Null Energy Condition (NEC):  $G_{\mu\nu}\eta^\mu\eta^\nu \geq 0$  for any null vector  $\eta^\mu$ ,  
 $G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu}$ ; for minimal coupling to gravity:  $T_{\mu\nu}\eta^\mu\eta^\nu \geq 0$ .
- **Trapped surface  $\Rightarrow$  Singularity.**



## Galileons and Horndeski

### «Galileon»

A. Nicolis, R. Rattazzi and E. Trincherini, Phys. Rev. D **79**, 064036 (2009)  
doi:10.1103/PhysRevD.79.064036, arXiv:0811.2197 [hep-th].

- Invariance under  $\partial^\mu \pi \rightarrow \partial^\mu \pi + b^\mu$ , where  $b^\mu$  is a constant vector.
- Second order field equations despite the second order derivatives in the Lagrangian.

### First article

G. W. Horndeski, Int. J. Theor. Phys. **10**, 363 (1974)  
doi:10.1007/BF01807638.



## Horndesky theory

$$S = \int d^4x (\mathcal{L}_2 + \mathcal{L}_3 + \mathcal{L}_4 + \mathcal{L}_5)$$

$$\mathcal{L}_2 = K(\phi, X),$$

$$\mathcal{L}_3 = -G_3(\phi, X)\square\phi,$$

$$\mathcal{L}_4 = G_4(\phi, X)R + G_{4X} [(\square\phi)^2 - (\nabla_\mu \nabla_\nu \phi)^2],$$

$$\mathcal{L}_5 = G_5(\phi, X)G_{\mu\nu}\nabla^\mu\nabla^\nu\phi - \frac{1}{6}G_{5X} [(\square\phi)^3 - 3\square\phi(\nabla_\mu \nabla_\nu \phi)^2 + 2(\nabla_\mu \nabla_\nu \phi)^3].$$

- $X = -\frac{1}{2}\nabla_\mu\phi\nabla^\mu\phi$ ,  $\square\phi = \nabla_\mu\nabla^\mu\phi$ ;
- $R$  – scalar curvature;
- $G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R$  – Einstein tensor;
- metric signature  $(-, +, +, +)$ .

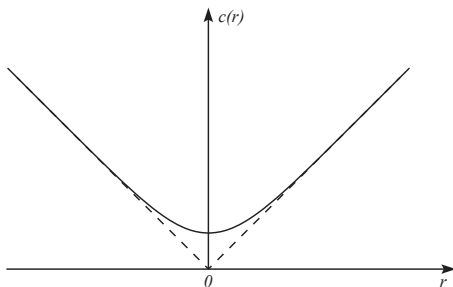


# Metric

## General form

$$ds^2 = -a^2(r)dt^2 + \frac{dr^2}{b^2(r)} + c^2(r)(d\theta^2 + \sin^2\theta d\varphi^2)$$

- Gauge  $a(r) = b(r)$ .
- Asymptotic behaviour:  $a(r) \rightarrow 1$ ,  $c(r) \rightarrow |r|$  as  $r \rightarrow \pm\infty$ .
- $c(r)$  behaviour:





## Spherical harmonics

$$s(t, r, \theta, \varphi) = \sum_{l=0}^{\infty} \sum_{m=-l}^l s_{lm}(t, r) Y_l^m(\theta, \varphi),$$

$$V_a(t, r, \theta, \varphi) = \overset{\gamma}{\nabla}_a \Phi_1(t, r, \theta, \varphi) + E_a^b \overset{\gamma}{\nabla}_b \Phi_2(t, r, \theta, \varphi),$$

$$T_{ab}(t, r, \theta, \varphi) = \overset{\gamma}{\nabla}_a \overset{\gamma}{\nabla}_b \Psi_1 + \gamma_{ab} \Psi_2(t, r, \theta, \varphi) + \frac{1}{2} \left( E_a^c \overset{\gamma}{\nabla}_c \overset{\gamma}{\nabla}_b \Psi_3(t, r, \theta, \varphi) + E_b^c \overset{\gamma}{\nabla}_c \overset{\gamma}{\nabla}_a \Psi_3(t, r, \theta, \varphi) \right).$$

- $\gamma_{ab}$  is  $S^2$  metric;
- $\overset{\gamma}{\nabla}_a$  is covariant derivative for  $\gamma_{ab}$ ;
- $E_{ab} = \sqrt{\det \gamma} \varepsilon_{ab}$ ,  $\varepsilon_{ab}$  – Levi-Civita symbol,  $\varepsilon_{\theta\varphi} = 1$ ;
- $\Phi_1, \Phi_2, \Psi_1, \Psi_2, \Psi_3$  are scalar functions;
- $Y_l^m(\theta, \varphi)$  are spherical harmonics.



## Perturbations (terms with $E_{ab}$ )

$$\delta\phi = 0, \quad h_{tt} = 0, \quad h_{tr} = 0, \quad h_{rr} = 0,$$

$$h_{ta} = \sum_{l=0}^{\infty} \sum_{m=-l}^l h_{0,lm}(t, r) E_{ab} \partial^b Y_l^m(\theta, \varphi),$$

$$h_{ra} = \sum_{l=1}^{\infty} \sum_{m=-l}^l h_{1,lm}(t, r) E_{ab} \partial^b Y_l^m(\theta, \varphi),$$

$$h_{ab} = \frac{1}{2} \sum_{l=2}^{\infty} \sum_{m=-l}^l h_{2,lm}(t, r) \left[ E_a^c \overset{\gamma}{\nabla}_c \overset{\gamma}{\nabla}_b Y_l^m(\theta, \varphi) + E_b^c \overset{\gamma}{\nabla}_c \overset{\gamma}{\nabla}_a Y_l^m(\theta, \varphi) \right].$$

Interval covariance under  $x^\mu \rightarrow x^\mu + \xi^\mu$

$$\xi_t = 0, \quad \xi_r = 0, \quad \xi_a = \sum_{l=0}^{\infty} \sum_{m=-l}^l \Lambda_{lm}(t, r) E_{ab} \partial^b Y_l^m(\theta, \varphi).$$





## Interval covariance under $x^\mu \rightarrow x^\mu + \xi^\mu$

$$h_{\mu\nu} \rightarrow h_{\mu\nu} + \nabla_\mu \xi_\nu + \nabla_\nu \xi_\mu$$

- $\xi_t = 0$ ,  $\xi_r = 0$ ,  $\xi_a = \sum_{l=0}^{\infty} \sum_{m=-l}^l \Lambda_{lm}(t, r) E_{ab} \partial^b Y_l^m(\theta, \varphi)$ ;
- $\nabla_\mu$  is the covariant derivative for  $g_{\mu\nu}$ .

$$h_{0,lm} \rightarrow h_{0,lm} + \dot{\Lambda}_{lm}(t, r),$$

$$h_{1,lm} \rightarrow h_{1,lm} + \Lambda'_{lm}(t, r) + 2\frac{c'}{c} \Lambda_{lm}(t, r),$$

$$h_{2,lm} \rightarrow h_{2,lm} + 2\Lambda_{lm}(t, r).$$

### Regge-Wheeler approach with $l \geq 2$

- $h_{2,lm} = 0$ ;
- $m = 0$  without loss of generality.



## Lagrangian for perturbations

$$\mathcal{L}^{(2)} = \frac{l(l+1)}{4(l-1)(l+2)} [A\dot{q}^2 - Bq'^2 - l(l+1)Cq^2 - V(r)q^2],$$

$$\text{where } A = \frac{c^2 \mathcal{H}^2}{a^2 \mathcal{G}}, \quad B = a^2 c^2 \frac{\mathcal{H}^2}{\mathcal{F}}, \quad C = a^2 \mathcal{H},$$

- $V(r)$  – effective potential;
- $C$ -part – wave propagation along the angular direction.



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### Stability conditions

T. Kobayashi, H. Motohashi and T. Suyama,  
 Phys. Rev. D **85**, 084025 (2012); **96**, 109903(E) (2017),  
 doi:10.1103/PhysRevD.96.109903, doi:10.1103/PhysRevD.85.084025,  
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### Stability conditions

- $\mathcal{F} = 2 \left( G_4 + \frac{a^2}{2} \phi' X' G_{5X} - X G_{5\phi} \right) > 0$  – radial direction,



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- $\mathcal{H} = 2 \left[ G_4 - 2X G_{4X} + X \left( a^2 \frac{c'}{c} \phi' G_{5X} + G_{5\phi} \right) \right] > 0$  – angular dir.



## Lagrangian for perturbations

$$\mathcal{L}^{(2)} = \frac{l(l+1)}{4(l-1)(l+2)} \left[ \mathcal{A} \dot{q}^2 - \mathcal{B} q'^2 - l(l+1) \mathcal{C} q^2 - V(r) q^2 \right],$$

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## Perturbations (terms without $E_{ab}$ )

$$\delta\phi = \sum_{l=0}^{\infty} \sum_{m=-l}^l \delta\phi_{lm}(t, r) Y_l^m(\theta, \varphi), \quad h_{tt} = a^2 \sum_{l=0}^{\infty} \sum_{m=-l}^l H_{0,lm}(t, r) Y_l^m(\theta, \varphi),$$

$$h_{tr} = \sum_{l=0}^{\infty} \sum_{m=-l}^l H_{1,lm}(t, r) Y_l^m(\theta, \varphi), \quad h_{rr} = \frac{1}{a^2} \sum_{l=0}^{\infty} \sum_{m=-l}^l H_{2,lm}(t, r) Y_l^m(\theta, \varphi),$$

$$h_{ta} = \sum_{l=0}^{\infty} \sum_{m=-l}^l \beta_{lm} \partial_a Y_l^m(\theta, \varphi), \quad h_{ra} = \sum_{l=0}^{\infty} \sum_{m=-l}^l \alpha_{lm} \partial_a Y_l^m(\theta, \varphi),$$

$$h_{ab} = c^2 \sum_{l=0}^{\infty} \sum_{m=-l}^l \left[ K_{lm}(t, r) \gamma_{ab} Y_l^m(\theta, \varphi) + G_{lm}(t, r) \nabla_a \nabla_b Y_l^m(\theta, \varphi) \right].$$





## Interval covariance under $x^\mu \rightarrow x^\mu + \xi^\mu$

$$h_{\mu\nu} \rightarrow h_{\mu\nu} + \nabla_\mu \xi_\nu + \nabla_\nu \xi_\mu$$

- $\nabla_\mu$  is the covariant derivative for  $g_{\mu\nu}$ .

### Gauge functions

$$\xi_0 = \sum_{l=0}^{\infty} \sum_{m=-l}^l T_{lm}(t, r) Y_l^m(\theta, \varphi),$$

$$\xi_r = \sum_{l=0}^{\infty} \sum_{m=-l}^l R_{lm}(t, r) Y_l^m(\theta, \varphi),$$

$$\xi_a = \sum_{l=0}^{\infty} \sum_{m=-l}^l \Theta_{lm}(t, r) \partial_a Y_l^m(\theta, \varphi).$$



## Gauge transformations of perturbations

$$H_{0,lm}(t, r) \rightarrow H_{0,lm}(t, r) + \frac{2}{a^2} \dot{T}_{lm}(t, r) - 2\frac{a'}{a} b^2 R_{lm}(t, r),$$

$$H_{1,lm}(t, r) \rightarrow H_{1,lm}(t, r) + \dot{R}(t, r) + T'(t, r) - 2\frac{a'}{a} T(t, r),$$

$$H_{2,lm}(t, r) \rightarrow H_{2,lm}(t, r) + 2b^2 R'_{lm}(t, r) - 2bb' R_{lm}(t, r),$$

$$\beta_{lm}(t, r) \rightarrow \beta_{lm}(t, r) + T_{lm}(t, r) + \dot{\Theta}_{lm}(t, r),$$

$$\alpha_{lm}(t, r) \rightarrow \alpha_{lm}(t, r) + R_{lm}(t, r) + \Theta'_{lm}(t, r) - 2\frac{c'}{c} \Theta_{lm}(t, r),$$

$$K_{lm}(t, r) \rightarrow K_{lm}(t, r) + 2b^2 \frac{c'}{c} R_{lm}(t, r),$$

$$G_{lm}(t, r) \rightarrow G_{lm}(t, r) + \frac{2}{c^2} \Theta_{lm}(t, r).$$

- Gauge fixing:  $\beta(t, r) = 0$ ,  $K(t, r) = 0$  and  $G(t, r) = 0$ .



## Lagrangian for perturbations

$$\mathcal{L} = \frac{1}{2}\mathcal{K}_{ij}\dot{v}^i\dot{v}^j - \frac{1}{2}\mathcal{G}_{ij}v^{i'}v^{j'} - \frac{1}{2}\mathcal{Q}_{ij}v^i v^{j'} - \frac{1}{2}\mathcal{M}_{ij}v^i v^j,$$

- $i = \overline{1,2}, j = \overline{1,2}$ ;
- $v^1 \equiv \psi, v^2 \equiv \delta\phi$ ;
- $H_0 = -\frac{2}{2cc'\mathcal{H} + \Xi\phi'} \left( \frac{1}{a^2}\psi - c'\Xi\delta\phi' - l(l+1)c'\mathcal{H}\alpha \right)$ ;
- $\Xi = 2c^2 \left[ -XG_{3X} + 2a^2\frac{c'}{c}\phi' \{ G_{4X} + 2XG_{4XX} - (XG_{5\phi})_X \} \right. \\ \left. + G_{4\phi} + 2XG_{4\phi X} - \frac{1}{c^2}XG_{5X} + a^2\frac{c'^2}{c^2}(3XG_{5X} + 2X^2G_{5XX}) \right].$



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### Stability conditions

T. Kobayashi, H. Motohashi and T. Suyama, Phys. Rev. D **89**, 084042 (2014)  
doi:10.1103/PhysRevD.89.084042 arXiv:1402.6740 [gr-qc]

- $\mathcal{K}_{11} > 0, \det \mathcal{K} > 0$  – no ghosts.



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## Stability condition $\det \mathcal{K} > 0$

$$\det \mathcal{K} = \frac{4(l-1)(l+2)(2cc'\mathcal{H} + \Xi\phi')^2 \mathcal{F}(2\mathcal{P}_1 - \mathcal{F})}{l(l+1)a^4\mathcal{H}^2\phi'^2(2cc'\mathcal{H} + \Xi\phi')^2} > 0,$$

$$\text{where } \mathcal{P}_1 = \frac{(2cc'\mathcal{H} + \Xi\phi')}{2c^2\mathcal{H}^2} \cdot \frac{d}{dr} \left[ \frac{c^4\mathcal{H}^4}{(2cc'\mathcal{H} + \Xi\phi')^2} \right].$$



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$$2\mathcal{P}_1 - \mathcal{F} > 0$$



## No-go theorem

### Stability conditions

- odd sector:  $\mathcal{F} > 0, \mathcal{H} > 0$ ;
- even sector:  $2\mathcal{P}_1 - \mathcal{F} > 0$ .

$$Q = \frac{2cc'\mathcal{H} + \Xi\phi'}{c^2\mathcal{H}^2} \Rightarrow 2\mathcal{P}_1 - \mathcal{F} = -2\frac{Q'}{Q^2} - \mathcal{F} > 0,$$

$$\frac{Q'}{Q^2} < -\frac{1}{2}\mathcal{F} \xrightarrow{\int \text{from } r \text{ to } r' > r} Q^{-1}(r) - Q^{-1}(r') < -\frac{1}{2} \int_r^{r'} \mathcal{F} dr.$$





## No-go theorem

- $Q^{-1}(r) < 0$  at some  $r$ :

$$Q^{-1}(r') > Q^{-1}(r) + \frac{1}{2} \int_r^{r'} \mathcal{F} dr.$$

- $Q^{-1}(r') > 0$  at some  $r'$ :

$$Q^{-1}(r) < Q^{-1}(r') - \frac{1}{2} \int_r^{r'} \mathcal{F} dr.$$

If  $\int_r^{r'} \mathcal{F} dr$  diverges at  $r \rightarrow \pm\infty$ ,  $Q$  has to be singular.



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general relativity restores away from the wormhole

$$\begin{cases} G_4 \rightarrow M_{Pl}^2/2 \\ G_5 \rightarrow 0 \end{cases} \text{ at } r \rightarrow \pm\infty \Rightarrow \mathcal{F} \rightarrow M_{Pl}^2.$$

Singular  $Q$ 

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  - if  $\mathcal{H}$  is singular  $Q \rightarrow 0$  not to  $\infty$
- }  $\Rightarrow$   $\mathbf{1}$  is not singular;

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- $$\left. \vphantom{\begin{matrix} \bullet c \text{ is bounded from below} \\ \bullet \mathcal{H} > 0 \end{matrix}} \right\} \Rightarrow \text{denominator is not equal to zero;}$$
- 
- $c'$  is not singular
  - if  $\mathcal{H}$  is singular  $Q \rightarrow 0$  not to  $\infty$
- $$\left. \vphantom{\begin{matrix} \bullet c' \text{ is not singular} \\ \bullet \text{if } \mathcal{H} \text{ is singular } Q \rightarrow 0 \text{ not to } \infty \end{matrix}} \right\} \Rightarrow \mathbf{1} \text{ is not singular;}$$
- 
- $\Xi$  is singular
  - $\phi'$  is singular
- $$\left. \vphantom{\begin{matrix} \bullet \Xi \text{ is singular} \\ \bullet \phi' \text{ is singular} \end{matrix}} \right\} \underline{\text{the Lagrangian is singular.}}$$



## Conclusion

- The analysis of both even and odd sector was made for  $l \geq 2$  within the Regge-Wheler approach.
- Stability conditions were obtained. They coincide with arXiv:1202.4893 [gr-qc] and arXiv:1402.6740 [gr-qc].
- No-go theorem proved: there are no static spherically symmetric Lorentzian wormholes in Horndeski theory.

*Thank you for your attention!*

